# **Ray Tracing**

Brian Curless CSE 557 Autumn 2017

#### Reading

#### Required:

- Marschner and Shirley, Ch. 4, Section 13.1-13.2 (online handout)
- Triangle intersection (online handout)

#### Further reading:

- Shirley errata on syllabus page, needed if you work from his book instead of the handout, which has already been corrected.
- ◆ T. Whitted. An improved illumination model for shaded display. Communications of the ACM 23(6), 343-349, 1980.
- A. Glassner. An Introduction to Ray Tracing. Academic Press, 1989.
- K. Turkowski, "Properties of Surface Normal Transformations," Graphics Gems, 1990, pp. 539-547.

#### **Geometric optics**

Modern theories of light treat it as both a wave and a particle.

We will take a combined and somewhat simpler view of light – the view of **geometric optics**.

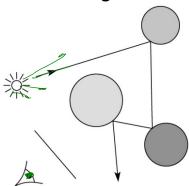
Here are the rules of geometric optics:

- Light is a flow of photons with wavelengths.
   We'll call these flows "light rays."
- Light rays travel in straight lines in free space.
- Light rays do not interfere with each other as they cross.
- Light rays obey the laws of reflection and refraction.
- Light rays travel from the light sources to the eye, but the physics is invariant under path reversal (reciprocity).

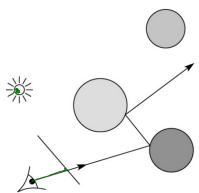
## Eye vs. light ray tracing

Where does light begin?

At the light: light ray tracing (a.k.a., forward ray tracing or photon tracing)



At the eye: eye ray tracing (a.k.a., backward ray tracing)

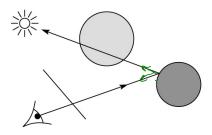


We will generally follow rays from the eye into the scene.

# **Precursors to ray tracing**

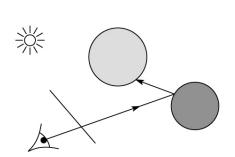
#### Local illumination

 Cast one eye ray, then shade according to light



#### Appel (1968)

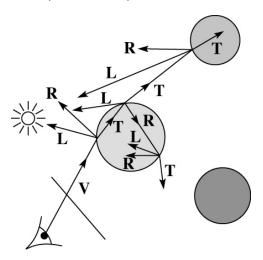
• Cast one eye ray + one ray to light



## Whitted ray-tracing algorithm

In 1980, Turner Whitted introduced ray tracing to the graphics community.

- ◆ Combines eye ray tracing + rays to light
- Recursively traces rays

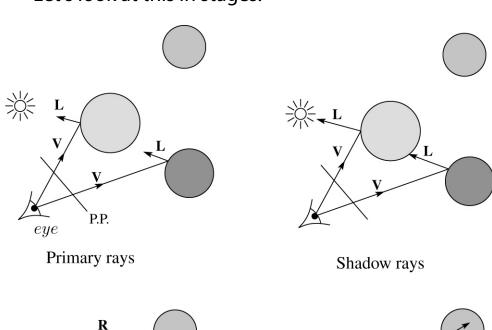


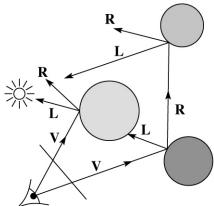
#### Algorithm:

- 1. For each pixel, trace a **primary ray** in direction **V** to the first visible surface.
- 2. For each intersection, trace **secondary rays**:
  - ullet Shadow rays in directions  $L_i$  to light sources
  - **Reflected ray** in direction **R**.
  - Refracted ray or transmitted ray in direction T.

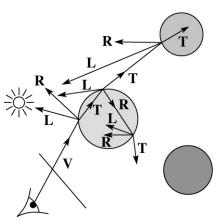
# Whitted algorithm (cont'd)

Let's look at this in stages:





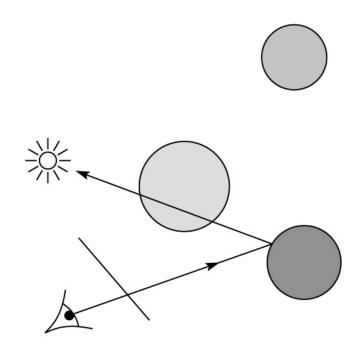
Reflection rays

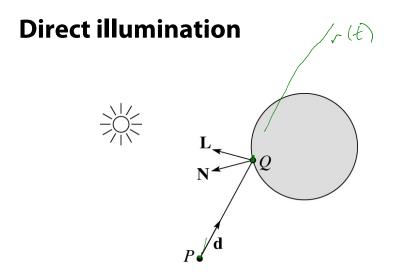


Refracted rays

# Ray casting and local illumination

Now let's actually build the ray tracer in stages. We'll start with ray casting and local illumination:





A ray is defined by an origin P and a unit direction  ${\bf d}$  and is parameterized by t>0:

$$\mathbf{r}(t) = P + t \, \mathbf{d}$$

Let  $I(P, \mathbf{d})$  be the intensity seen along a ray. Then:

$$I(P, \mathbf{d}) = I_{\text{direct}}$$

where

ullet  $I_{
m direct}$  is computed from the Blinn-Phong model

#### **Shading in "Trace"**

The Trace project uses a version of the Blinn-Phong shading equation we derived in class, with three modifications:

Distance attenuation is clamped to be at most 1:

$$A_j^{dist} = \min\left\{1, \frac{1}{a_j + b_j r_j + c_j r_j^2}\right\}$$

- Shadow attenuation  $A^{\text{shadow}}$  is included.
- We will use  $k_d$  for the ambient reflection coefficient.

Here's what the shading equation should look like:

$$I = k_e + \sum_{j} \underbrace{k_d}_{La,j} + \underbrace{A_j^{shadow}}_{j} \underbrace{A_j^{dist}}_{L,j} I_{L,j} B_j \left[ k_d \left( \mathbf{N} \cdot \mathbf{L}_j \right) + k_s \left( \mathbf{N} \cdot \mathbf{H}_j \right)_+^{n_s} \right]$$

This is the shading equation to use in the Trace project!

## **Ray-tracing pseudocode**

We build a ray traced image by casting rays through each of the pixels.

#### function tracelmage (scene):

**for each** pixel (i, j) in image

$$A = pixelToWorld(i, j)$$

$$\underline{P} = \mathbf{C}$$

$$\mathbf{d} = (A - P) / ||A - P||$$

I(i,j) = traceRay(scene, P, d)

end for

#### end function

ction RADA

**function** *traceRay* (scene, *P*, **d**):

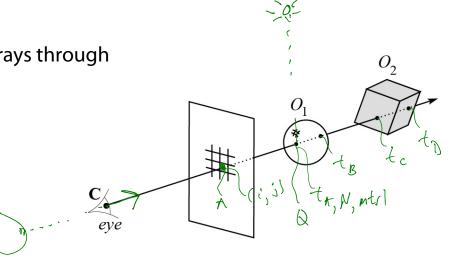
 $(t_{\cap}, \mathbf{N}, \mathsf{mtrl}) \leftarrow \mathsf{scene}.intersect(P, \mathbf{d})$ 

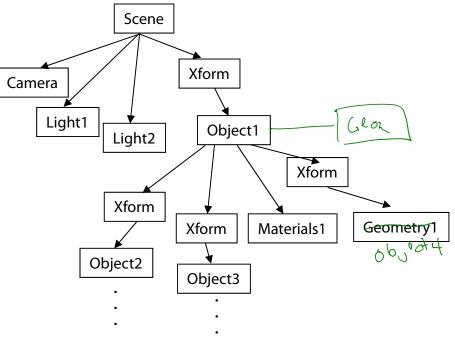
 $\widetilde{Q} \leftarrow \text{ray}(P, \mathbf{d}) \text{ evaluated at } t$ 

I = shade (wtr), N, scene, d, Q

return I

end function





#### **Shading pseudocode**

Next, we need to calculate the color returned by the *shade* function.

```
function shade (mtrl, scene, Q, N, d):

I \leftarrow \text{mtrl}.k_e + \text{mtrl}.k_a + \text{land}

for each light source Light in scene do:

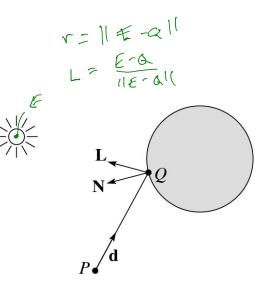
atten = \text{Light} -> distanceAttenuation}(Q)

L = \text{Light} -> \text{getDirection}(Q)

I \leftarrow I + \text{atten*(diffuse term + specular term)}

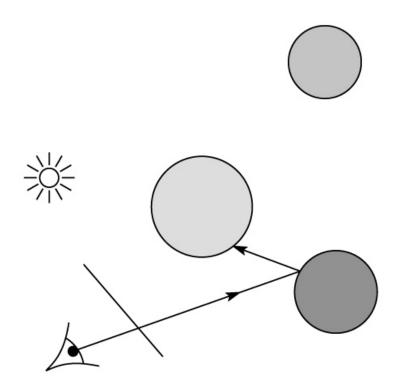
end for

man k_d = mbient + mbient + man k_d = mbient + m
```



# Ray casting with shadows

Now we'll add shadows by casting shadow rays:



#### **Shading with shadows**

To include shadows, we need to modify the shade function:

```
function shade (mtrl, scene, Q, N, d):

I \leftarrow \mathsf{mtrl}.k_e + \mathsf{mtrl}.k_d * I_{La}
for each light source Light in scene do:
\mathsf{atten} = \mathsf{Light} -> \mathit{distanceAttenuation}(Q) *
\mathsf{Light} -> \mathit{shadowAttenuation}(Q) *
\mathsf{L} = \mathsf{Light} -> \mathsf{getDirection}(Q)
I \leftarrow I + \mathsf{vatten*}(\mathsf{diffuse\ term} + \mathsf{specular\ term})
end for return I
end function
```

#### **Shadow attenuation**

Computing a shadow can be as simple as checking to see if a ray makes it to the light source.

For a point light source:

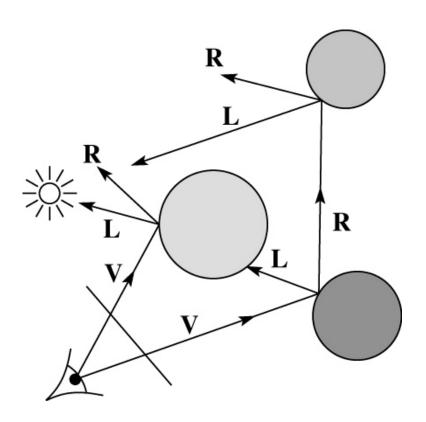
```
function PointLight :: shadowAttenuation (scene, Q)
L = \text{getDirection}(Q)
(t_{\bigcirc}, \mathbf{N}, \text{mtrl}) \leftarrow \text{scene.} intersect(Q, \mathbf{L})
Compute \ t_{\text{light}}
\mathbf{if} \ (t_{\bigcirc} < t_{\text{light}}) \ \mathbf{then:}
\mathbf{atten} = (0, 0, 0)
\mathbf{else}
\mathbf{atten} = (1, 1, 1)
\mathbf{end} \ \mathbf{if}
\mathbf{return} \ \mathbf{atten}
\mathbf{end} \ \mathbf{function}
```

Note: we will later handle color-filtered shadowing, so this function needs to return a *color* value.

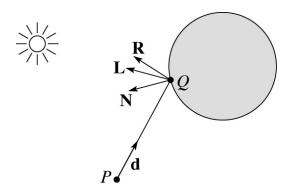
For a directional light,  $t_{\text{light}} = \infty$ .

# **Recursive ray tracing with reflection**

Now we'll add reflection:



## **Shading with reflection**



Let  $I(P, \mathbf{d})$  be the intensity seen along a ray. Then:

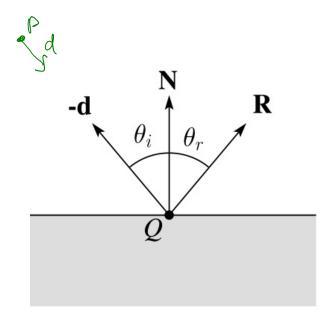
$$I(P, \mathbf{d}) = I_{\text{direct}} + I_{\text{reflected}}$$

where

- ullet  $I_{
  m direct}$  is computed from the Blinn-Phong model, plus shadow attenuation
- $I_{\text{reflected}} = k_r I(Q, \mathbf{R})$

Typically, we set  $k_r = k_s$ . ( $k_r$  is a color value.)

## Reflection



Law of reflection:

$$\theta_i = \theta_r$$

 $\boldsymbol{R}$  is co-planar with  $\boldsymbol{d}$  and  $\boldsymbol{N}.$ 

## Ray-tracing pseudocode, revisited

```
function traceRay (scene, P, d):
(t_{\bigcirc}, \mathbf{N}, \mathsf{mtrl}) \leftarrow \mathsf{scene}.intersect(P, \mathbf{d})
Q \leftarrow \mathsf{ray}(P, \mathbf{d}) \; \mathsf{evaluated} \; \mathsf{at} \; t
I = \mathit{shade} \; (\mathsf{scene}, \mathsf{mtrl}, Q, \mathbf{N}, \mathbf{d})
\mathbf{R} = \mathit{reflectDirection}(-\phi_{\gamma} \mathcal{N}) \quad )
I \leftarrow I + \mathsf{mtrl}.k_r * \mathit{traceRay}(\mathsf{scene}, Q, \mathbf{R})
\mathsf{return} \; I
\mathsf{end} \; \mathsf{function}
```

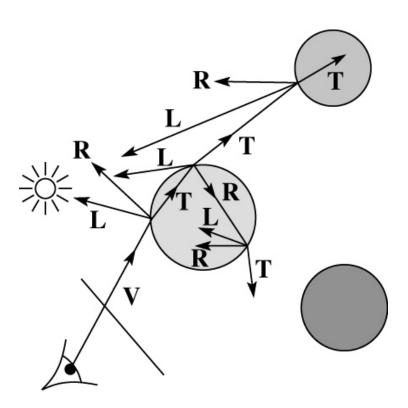
#### **Terminating recursion**

**Q**: How do you bottom out of recursive ray tracing?

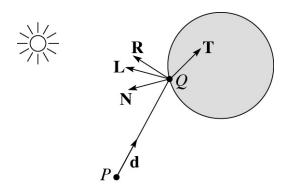
Possibilities:

# Whitted ray tracing

Finally, we'll add refraction, giving us the Whitted ray tracing model:



#### **Shading with reflection and refraction**



Let  $I(P, \mathbf{d})$  be the intensity seen along a ray. Then:

$$I(P, \mathbf{d}) = I_{\text{direct}} + I_{\text{reflected}} + I_{\text{transmitted}}$$

where

- ullet  $I_{
  m direct}$  is computed from the Blinn-Phong model, plus shadow attenuation
- $I_{\text{reflected}} = k_r I(Q, \mathbf{R})$
- $I_{\text{transmitted}} = k_t I(Q, \mathbf{T})$

Typically, we set  $k_r = k_{\rm s}$  and  $k_t = 1 - k_{\rm s}$  (or (0,0,0), if opaque, where  $k_t$  is a color value).

[Generally,  $k_r$  and  $k_t$  are determined by "Fresnel reflection," which depends on angle of incidence and changes the polarization of the light. This is discussed in Marschner's textbook and can be implemented for extra credit.]

## Refraction

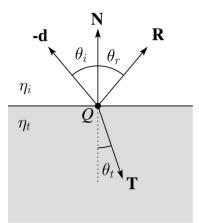


Snell's law of refraction:

$$\eta_i \sin \theta_i = \eta_t \sin \theta_t$$

where  $\eta_i$ ,  $\eta_t$  are indices of refraction.

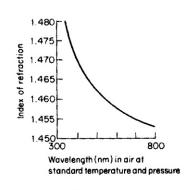
In all cases, **R** and **T** are co-planar with **d** and **N**.



The index of refraction is material dependent.

It can also vary with wavelength, an effect called **dispersion** that explains the colorful light rainbows from prisms. (We will generally assume no dispersion.)

Medium	Index of refraction
Vaccum Air Water Fused quartz Glass, crown Glass, dense flint Diamond	1 1.0003 1.33 1.46 1.52 1.66 2.42



Index of refraction variation for fused quartz

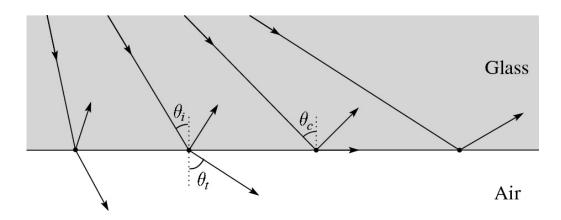
#### **Total Internal Reflection**

The equation for the angle of refraction can be computed from Snell's law:  $M_{i} sin b_{i} = M_{t} sin b_{t}$   $\theta_{t} = sin \left( \frac{M_{i}}{M_{t}} sin b_{i} \right)$ 

What "bad thing" can happen when  $\eta_i > \eta_t$ ?

When  $\theta_i$  is exactly 90°, we say that  $\theta_i$  has achieved the "critical angle"  $\theta_c$ .

For  $\theta_i > \theta_c$ , no rays are transmitted, and only reflection occurs, a phenomenon known as "total internal reflection" or TIR.



## Marschner's notation

Marschner uses different symbols. Here is the translation between them:

$$\mathbf{r} = \mathbf{R}$$

$$\mathbf{t} = \mathbf{T}$$

$$\phi = \theta_t$$

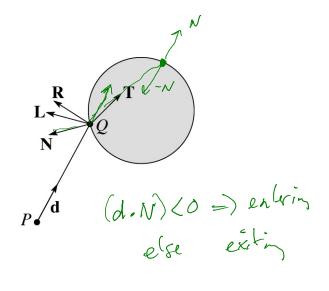
$$\theta = \theta_r = \theta_i$$

$$n = \eta_i$$

$$n_t = \eta_t$$

#### Ray-tracing pseudocode, revisited

```
function traceRay (scene, P, d):
     (t_{\cap}, \mathbf{N}, \mathsf{mtrl}) \leftarrow \mathsf{scene}.intersect(P, \mathbf{d})
      Q \leftarrow \text{ray}(P, \mathbf{d}) \text{ evaluated at } t
     I = shade (scene, mtrl, Q, N, d)
     \mathbf{R} = reflectDirection(\mathbf{N}, \mathbf{d})
     I \leftarrow I + \text{mtrl.} k_r * traceRay (scene, Q, \mathbf{R})
      if ray is entering object then
            \eta_i = \text{index\_of\_air} (=1.0003)
            \eta_t = \text{mtrl.index}
      else
            \eta_i = mtrl.index
            \eta_t = \text{index\_of\_air} (=1.0003)
     if (notTIR(d, N, M;, Mt
                                                              )) then
            T = refractDirection(
           I \leftarrow I + \text{mtrl.} k_t * traceRay (scene, Q, T)
      end if
      return I
end function
```



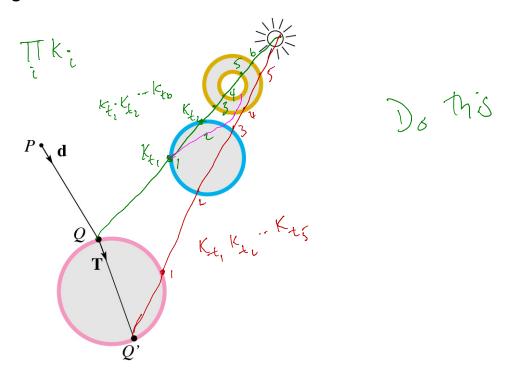
**Q**: How do we decide if a ray is entering the object?

## Terminating recursion, incl. refraction

**Q**: *Now* how do you bottom out of recursive ray tracing?

#### **Shadow attenuation (cont'd)**

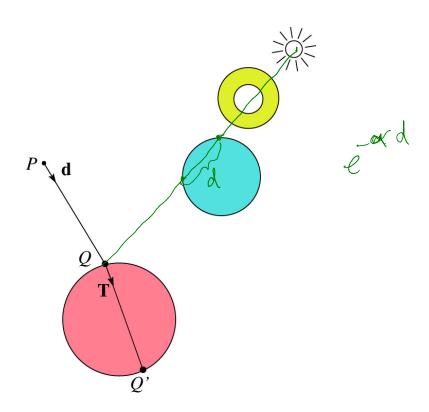
**Q**: What if there are transparent objects along a path to the light source?



We'll take the view that the color is really only at the surface, like a glass object with a colored transparency coating on it. In this case, we multiply in the transparency constant,  $k_t$ , every time an object is entered or exited, possibly more than once for the same object.

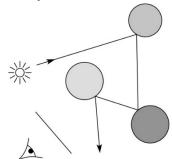
#### **Shadow attenuation (cont'd)**

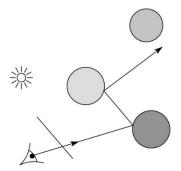
Another model would be to treat the glass as solidly colored in the interior. Marschner's textbook describes a the resulting volumetric attenuation based on Beer's Law, which you can implement for extra credit.



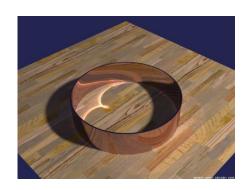
# **Photon mapping**

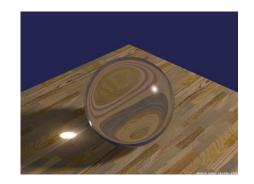
Combine light ray tracing (photon tracing) and eye ray tracing:





...to get **photon mapping**.







Renderings by Henrik Wann Jensen:

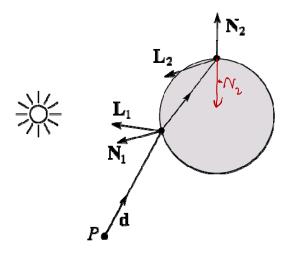
http://graphics.ucsd.edu/~henrik/images/caustics.html

# Normals and shading, reflection, and refraction when inside

When a ray is inside an object and intersects the object's surface on the way out, the normal will be pointing **away** from the ray (i.e., the normal always points to the outside by default).

You must *negate* the normal before doing any of the **shading**, **reflection**, **and refraction** that follows.

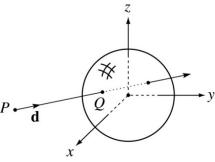
Finally, when shading a point inside of an object, apply  $k_t$  to the ambient component, since that "ambient light" had to pass through the object to get there in the first place.



#### **Intersecting rays with spheres**

Now we've done everything except figure out what that "scene. *intersect* (*P*, **d**)" function does.

Mostly, it calls each object to find out the *t*-value at which the ray intersects the object. Let's start with intersecting spheres...



#### Given:

• The coordinates of a point along a ray passing through *P* in the direction **d** are:

$$x = P_x + td_x$$

$$y = P_y + td_y$$

$$z = P_z + td_z$$

**Find**: The *t* at which the ray intersects *S*.

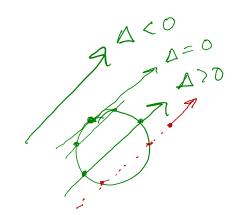
## **Intersecting rays with spheres**

#### **Solution by substitution**:

$$x^{2} + y^{2} + z^{2} - 1 = 0$$

$$(P_{x} + td_{x})^{2} + (P_{y} + td_{y})^{2} + (P_{z} + td_{z})^{2} - 1 = 0$$

$$at^{2} + bt + c = 0$$



where

$$a = d_x^2 + d_y^2 + d_z^2$$

$$b = 2(P_x d_x + P_y d_y + P_z d_z)$$

$$c = P_x^2 + P_y^2 + P_z^2 - 1$$

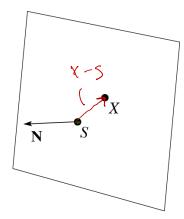
**Q**: What are the solutions of the quadratic equation in *t* and what do they mean?

and what do they mean?
$$t = -\frac{b}{2a} + 1 - \sqrt{b^2 - 4ac}$$

$$\Delta = b^2 - 4ac$$

**Q**: What is the normal to the sphere at a point (x, y, z) on the sphere?  $\bigvee = \begin{bmatrix} \times \\ \gamma \\ \emptyset \end{bmatrix}$ 

## **Ray-plane intersection**



Next, we will considering intersecting a ray with a plane.

To do this, we first need to define the plane equation.

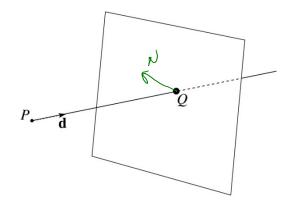
Given a point S on a plane with normal  $\mathbb{N}$ , how would we determine if another point X is on the plane?

(Hint: start by forming the vector X - S.)

$$\lambda \cdot (x-S) = 0$$

This is the plane equation!

#### Ray-plane intersection (cont'd)



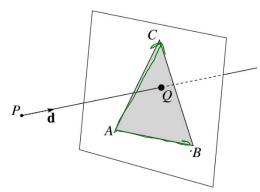
Now consider a ray intersecting a plane, where the plane equation is:

$$\mathbf{N} \cdot X = k$$

We can solve for the intersection parameter (and thus the point) by substituting X with the ray P + t  $\mathbf{d}$ ;

$$N \cdot (P+td) = k$$
 $N \cdot P + t N \cdot d = k$ 
 $t = \frac{k - N \cdot P}{N \cdot d}$ 
 $N \cdot d = 0 \Rightarrow d // Plane$ 

#### **Ray-triangle intersection**



To intersect with a triangle, we first solve for the equation of its supporting plane.

How might we compute the (un-normalized) normal?

$$N \approx (B-A) \times (C-A)$$
  $\hat{N} = \frac{N}{1|N|}$ 

Given this normal, how would we compute *k*?

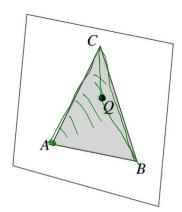
Using these coefficients, we can intersect the ray with the triangle to solve for Q.

Now, we need to decide if Q is inside or outside of the triangle...

#### **Computing barycentric coordinates**

Recall that given a point Q that is inside of triangle ABC, we can solve for Q's barycentric coordinates in a simple way:

$$\alpha = \frac{\text{SArea}(QBC)}{\text{SArea}(ABC)}$$
  $\beta = \frac{\text{SArea}(AQC)}{\text{SArea}(ABC)}$   $\gamma = \frac{\text{SArea}(ABQ)}{\text{SArea}(ABC)}$ 



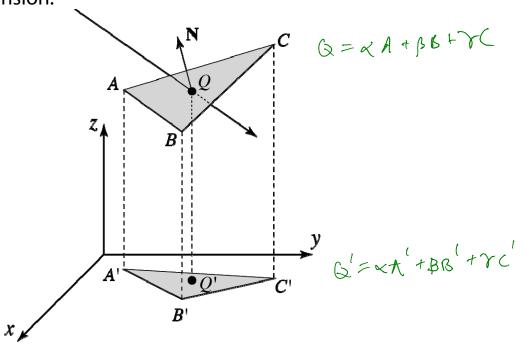
a AtBB+TC

mxv= ||u|||(v|| size N

Before, we computed signed areas with 2D cross products. How can we compute signed area with 3D cross products?  $\lambda_{cs} + \lambda_{cs} + \lambda_{cs$ 

## Improvement: project down to 2D first

Without loss of generality, we can make this determination after projecting down a dimension:



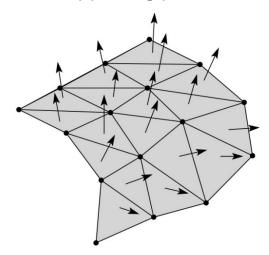
Why is this projection desirable? Faster—fower calculating in 21)

Which axis should you "project away"? (Hint: one of most dot product consider the triangle normal.)

## Per-face normals for triangle meshes

We will be shading and calculating reflections and refractions based on surface normals.

For a triangle mesh, we can make the natural assumption that each triangle has a constant normal (the normal of its supporting plane):



Recall the Blinn-Phong shading equation for a single light source (no ambient or emissive):

$$I = I_L B \left[ k_d (\mathbf{N} \cdot \mathbf{L}) + k_s (\mathbf{N} \cdot \mathbf{H})_+^{n_s} \right]$$

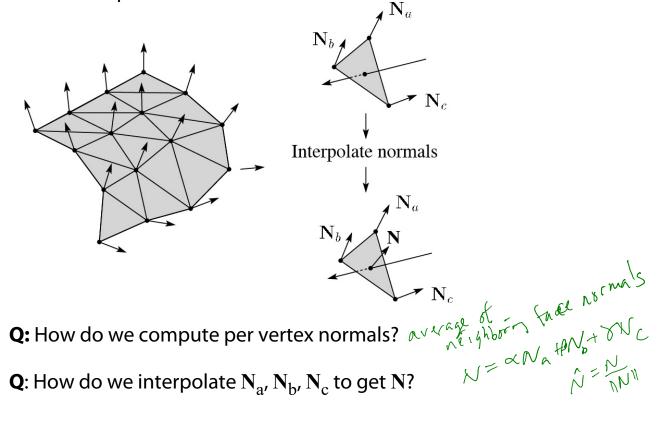
**Q**: If material properties  $(k_d, k_s, n_s)$  are constant over the mesh, how will shading vary within a triangle?

#### **Phong interpolated normals**

For better results, we can "hallucinate" smoothly varying within a triangle, a process called **Phong interpolation**. (Not the same as "Phong shading.")

#### Here's how it works:

- 1. Compute normals at the vertices.
- 2. Interpolate normals and normalize.
- 3. Shade, reflect, and refract using the interpolated normals.



#### **Interpolating vertex properties**

The barycentric coordinates can be used to interpolate a variety of vertex properties such as:

- material properties
- texture coordinates
- normals

For example:

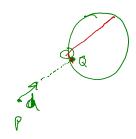
$$k_d(Q) = \alpha k_d(A) + \beta k_d(B) + \gamma k_d(C)$$

Interpolating normals, known as **Phong** *interpolation*, gives triangle meshes a smooth shading appearance. (Note: don't forget to normalize interpolated normals.)

## **Epsilons**

Due to finite precision arithmetic, we do not always get the exact intersection at a surface.

**Q**: What kinds of problems might this cause?



Q: How might we resolve this?  

$$+ > RAY \_ ERSILON \implies \cap$$
  
 $alse no \cap$ 

#### **Intersecting with xformed geometry**

In general, objects will be placed using transformations. What if the object being intersected were transformed by a matrix M?

Apply M<sup>-1</sup> to the ray first and intersect in object (local) coordinates! **Note**: *do not normalize* **d**'!

