Reading Required: • Section 2.7 • Chapter 4 • Section 13.1 - 13.2 Further reading: • T. Whitted. An improved illumination model for shaded display. Communications of the ACM 23(6), 343-349, 1980. • A. Glassner. An Introduction to Ray Tracing. Academic Press, 1989. • K. Turkowski, "Properties of Surface Normal Transformations," Graphics Gems, 1990, pp. 539-547.

3

Geometric optics

Modern theories of light treat it as both a wave and a particle.

We will take a combined and somewhat simpler view of light – the view of **geometric optics**.

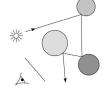
Here are the rules of geometric optics:

- Light is a flow of photons with wavelengths. We'll call these flows "light rays."
- Light rays travel in straight lines in free space.
- Light rays do not interfere with each other as they cross.
- Light rays obey the laws of reflection and refraction.
- Light rays travel from the light sources to the eye, but the physics is invariant under path reversal (reciprocity).

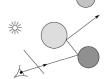
Eye vs. light ray tracing

Where does light begin?

At the light: light ray tracing (a.k.a., forward ray tracing or photon tracing)



At the eye: eye ray tracing (a.k.a., backward ray tracing)



We will generally follow rays from the eye into the scene.

Ray Tracing

Brian Curless CSE 557 Fall 2014

Precursors to ray tracing

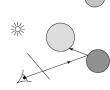
Local illumination

• Cast one eye ray, then shade according to light



Appel (1968)

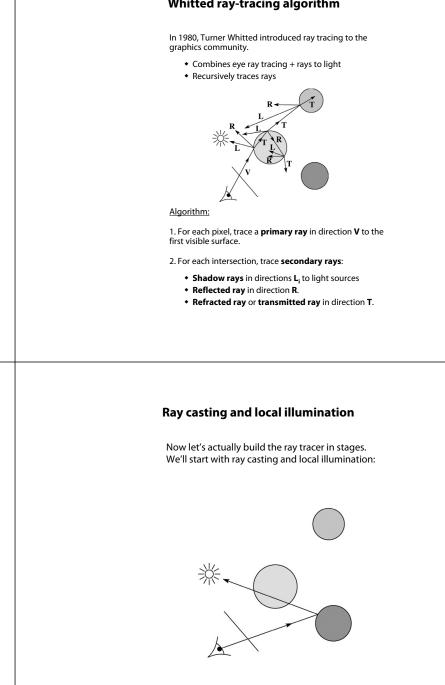
• Cast one eye ray + one ray to light



Whitted algorithm (cont'd)

Let's look at this in stages: 淡 eyePrimary rays Shadow rays Reflection rays Refracted rays

Whitted ray-tracing algorithm

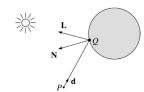


6

5

7

Direct illumination



A ray is defined by an origin P and a unit direction **d** and is parameterized by t > 0.

 $\mathbf{r}(t) = P + t\mathbf{d}$

Let *I*(*P*, **d**) be the intensity seen along a ray. Then:

 $I(P, \mathbf{d}) = I_{\text{direct}}$

where

• Idirect is computed from the Blinn-Phong model

Ray-tracing pseudocode

We build a ray traced image by casting rays through each of the pixels.

function traceImage (scene):

for each pixel (i,j) in image A = pixelToWorld(i,j) P = COP d = (A - P)/|| A - P|| I(i,j) = traceRay(scene, P, d)end for end function

function traceRay(scene, P, d):
 (t, N, mtrl) ← scene.intersect(P, d)
 Q ← ray (P, d) evaluated at t
 I = shade()
 return I
end function

Shading in "Trace"

The Trace project uses a version of the Blinn-Phong shading equation we derived in class, with two modifications:

• Distance attenuation is clamped to be at most 1:

$$A_j^{dist} = \min\left\{1, \frac{1}{a_j + b_j r_j + c_j r_j^2}\right\}$$

• Shadow attenuation A^{shadow} is included.

Here's what it should look like:

1

$$= k_e + k_s I_{Le} + \sum_j A_j^{shadow} A_j^{dist} I_{Lj} B_j \left[k_\sigma \left(\mathbf{N} \cdot \mathbf{L}_j \right) + k_s \left(\mathbf{N} \cdot \mathbf{H}_j \right)^{n_s} \right]$$

This is the shading equation to use in the Trace project!

Shading pseudocode

Next, we need to calculate the color returned by the *shade* function.

function shade(mtrl, scene, Q, N, d):

 $I \leftarrow mtrl.k_e + mtrl.k_a * I_{La}$ for each light source Light do: atten = Light -> distanceAttenuation(

)

 $L = Light \rightarrow getDirection ($)

 $I \leftarrow I + atten*(diffuse term + specular term)$

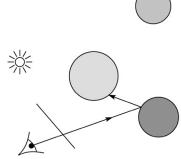
end for return |

end function

9

Ray casting with shadows

Now we'll add shadows by casting shadow rays:



Shading with shadows

To include shadows, we need to modify the shade function:

function shade(mtrl, scene, Q, N, d):

 $I \leftarrow mtrl.k_e + mtrl.k_a * I_{La}$ for each light source Light do: atten = Light -> distanceAttenuation(Q) * Light -> shadowAttenuation(L = Light -> getDirection (Q)

I ← I + atten*(diffuse term + specular term)

)

end for return |

end function

13

Shadow attenuation

Computing a shadow can be as simple as checking to see if a ray makes it to the light source. For a point light source:

function PointLight:shadowAttenuation(scene, P)

$$\label{eq:constraint} \begin{split} & \textbf{d} = \texttt{getDirection}(\mathsf{P}) \\ & (t, \textbf{N}, \mathsf{mtrl}) \leftarrow \texttt{scene.} \textit{intersect}(\textit{P}, \textbf{d}) \\ & \texttt{Compute} \ t_{light} \\ & \textbf{if} \ (t < t_{light}) \ \textbf{then:} \\ & \texttt{atten} = (0, 0, 0) \end{split}$$

else

atten = (1, 1, 1)

end if

return atten

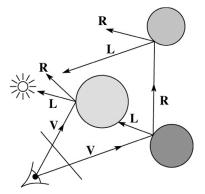
end function

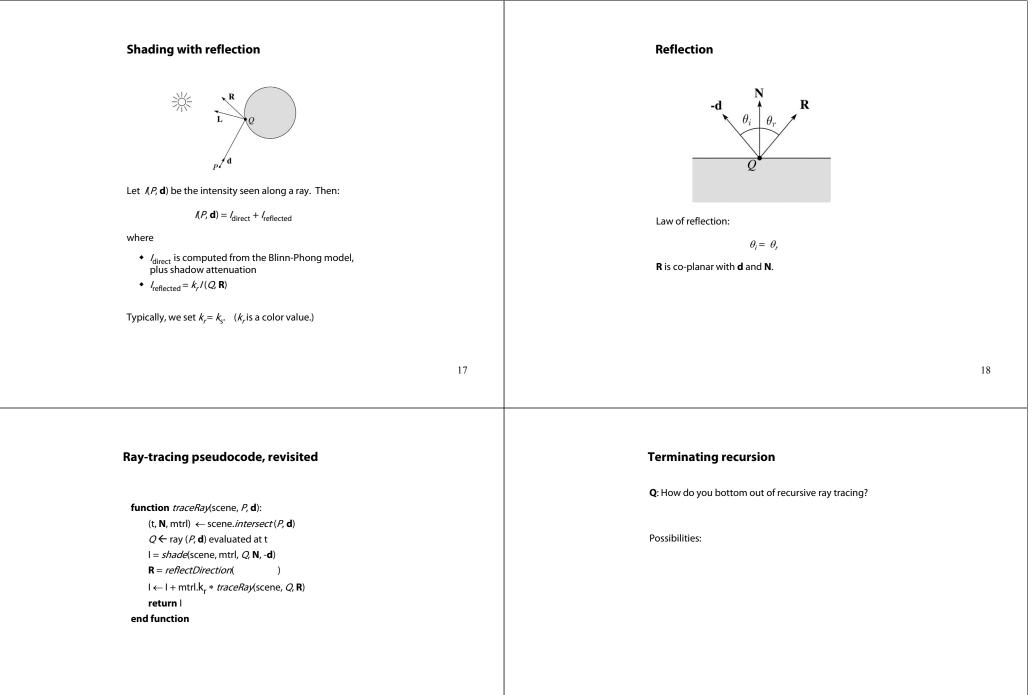
Note: we will later handle color-filtered shadowing, so this function needs to return a *color* value.

For a directional light, $t_{light} = \infty$.

Recursive ray tracing with reflection

Now we'll add reflection:

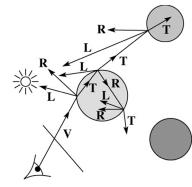




Whitted ray tracing

Shading with reflection and refraction

Finally, we'll add refraction, giving us the Whitted ray tracing model:



21

Refraction

Snell's law of refraction:

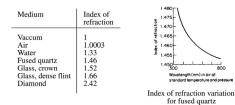
 $\eta_i \sin \theta_i = \eta_t \sin \theta_t$

where η_i , η_t are indices of refraction.

In all cases, **R** and **T** are co-planar with **d** and **N**.

The index of refraction is material dependent.

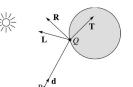
It can also vary with wavelength, an effect called **dispersion** that explains the colorful light rainbows from prisms. (We will generally assume no dispersion.)





T

23



Let *I*(*P*, **d**) be the intensity seen along a ray. Then:



where

Indirect is computed from the Blinn-Phong model, plus shadow attenuation

- $l_{\text{reflected}} = k_r / (Q, \mathbf{R})$
- $l_{\text{transmitted}} = k_t / (Q, \mathbf{T})$

Typically, we set $k_r = k_s$ and $k_t = 1 - k_s$ (or (0,0,0), if opaque, where k_r is a color value).

[Generally, k_r and k_r are determined by "Fresnel reflection," which depends on angle of incidence and changes the polarization of the light. This is discussed in Shirley's textbook and can be implemented for extra credit.]

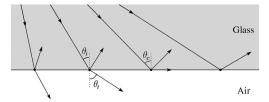
Total Internal Reflection

The equation for the angle of refraction can be computed from Snell's law:

What happens when $\eta_i > \eta_t$?

When θ_t is exactly 90°, we say that θ_i has achieved the "critical angle" θ_r .

For $\theta_i > \theta_c$, *no rays are transmitted*, and only reflection occurs, a phenomenon known as "total internal reflection" or TIR.



Shirley's notation

Shirley uses different symbols. Here is the translation between them:

 $\mathbf{r} = \mathbf{R}$ $\mathbf{t} = \mathbf{T}$ $\phi = \theta_t$ $\theta = \theta_r = \theta_i$ $n = \eta_i$ $n_t = \eta_t$

25

Terminating recursion, incl. refraction

Q: *Now* how do you bottom out of recursive ray tracing?

if ray is entering object then n_i = index_of_air n_t = mtrl.index else n_i = mtrl.index

Ray-tracing pseudocode, revisited

 $(t, \mathbf{N}, mtrl) \leftarrow scene.intersect(P, \mathbf{d})$

 $I \leftarrow I + mtrl.k_r * traceRay(scene, Q, R)$

 $Q \leftarrow ray (P, \mathbf{d})$ evaluated at t

I = shade(scene, mtrl, Q, N, -d)

R = *reflectDirection*(**N**, -**d**)

function *traceRay*(scene, *P*, **d**):

n_t = index_of_air if (*notTlR*()) then T = *refractDirection*() I ← I + mtrl.k_t * *traceRay*(scene, *Q*, T) end if return I end function

Q: How do we decide if a ray is entering the object?

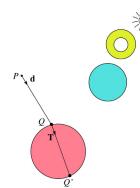
Shadow attenuation (cont'd)

Q: What if there are transparent objects along a path to the light source?

We'll take the view that the color is really only at the surface, like a glass object with a colored transparency coating on it. In this case, we multiply in the transparency constant, k_p every time an object is entered or exited, possibly more than once for the same object.

Shadow attenuation (cont'd)

Another model would be to treat the glass as solidly colored in the interior. Shirley's textbook describes a the resulting volumetric attenuation based on Beer's Law, which you can implement for extra credit.



29

31

Normals and shading when inside

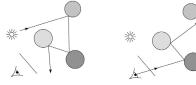
When a ray is inside an object and intersects the object's surface on the way out, the normal will be pointing *away* from the ray (i.e., the normal always points to the outside by default).

You must *negate* the normal before doing any of the shading, reflection, and refraction that follows.

Finally, when shading a point inside of an object, apply k_t to the ambient component, since that "ambient light" had to pass through the object to get there in the first place.

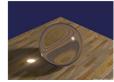
Photon mapping

Combine light ray tracing (photon tracing) and eye ray tracing:



...to get **photon mapping**.

Ø





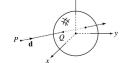
Renderings by Henrik Wann Jensen:

http://graphics.ucsd.edu/~henrik/ images/caustics.html

Intersecting rays with spheres

Now we've done everything except figure out what that "scene.intersect(P, d)" function does.

Mostly, it calls each object to find out the <u>t</u> value at which the ray intersects the object. Let's start with intersecting spheres...



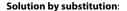
Given:

• The coordinates of a point along a ray passing through *P* in the direction **d** are:

$$x = P_x + td_x$$
$$y = P_y + td_y$$
$$z = P_z + td_z$$

• A unit sphere S centered at the origin defined by the equation:

Intersecting rays with spheres



$$x^{2} + y^{2} + z^{2} - 1 = 0$$

$$(P_{x} + td_{x})^{2} + (P_{y} + td_{y})^{2} + (P_{z} + td_{z})^{2} - 1 = 0$$

$$at^{2} + bt + c = 0$$

where

 $a = d_x^2 + d_y^2 + d_z^2$ $b = 2(P_x d_x + P_y d_y + P_z d_z)$ $c = P_x^2 + P_y^2 + P_z^2 - 1$

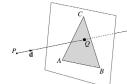
Q: What are the solutions of the quadratic equation in *t* and what do they mean?

Q: What is the normal to the sphere at a point (*x*,*y*,*z*) on the sphere?

33

This is the plane equation!

Ray-triangle intersection



To intersect with a triangle, we first solve for the equation of its supporting plane.

How might we compute the (un-normalized) normal?

Given this normal, how would we compute k?

Using these coefficients, we can solve for *Q*. Now, we need to decide if *Q* is inside or outside of the triangle.

Solution 1: Compute barycentric coordinates from 3D points.

What do you do with the barycentric coordinates?

Ray-plane intersection



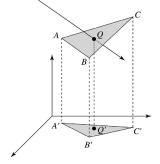
Next, we will considering intersecting a ray with a plane.

To do this, we first need to define the plane equation.

Given a point *S* on a plane with normal **N**, how would we determine if a point *X* is on the plane? Hint: start by forming the vector X - S.

2D inside-outside test

Without loss of generality, we can make this determination after projecting down a dimension:



If Q'is inside of A'B'C', then Q is inside of ABC.

Why is this projection desirable?

Which axis should you "project away"?

Interpolating vertex properties

The barycentric coordinates can also be used to interpolate vertex properties such as:

- material properties
- texture coordinates
- normals

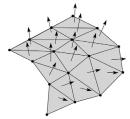
For example:

 $k_d(Q) = \alpha k_d(A) + \beta k_d(B) + \gamma k_d(C)$

Per-face normals for triangle meshes

We will be shading and calculating reflections and refractions based on surface normals.

For a triangle mesh, we can make the natural assumption that each triangle has a constant normal (the normal of its supporting plane):



Recall the Blinn-Phong shading equation for a single light source (no ambient or emissive):

 $I = I_{L}B\left[k_{d}(\mathbf{N}\cdot\mathbf{L}) + k_{s}(\mathbf{N}\cdot\mathbf{H})^{n_{s}}_{+}\right]$

Typically, **L** and **V** vary only a small amount over each triangle, if at all. Now assume constant material properties over the mesh. How will shading vary within a triangle?

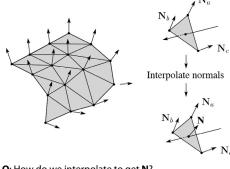
38

Phong interpolated normals

To get smoother results, we can "hallucinate" normals that vary smoothly within a triangle, a process called **Phong** interpolation. (Not the same as "Phong shading.")

Here's how it works:

- 1. Compute normals at the vertices.
- 2. Interpolate normals and normalize.
- 3. Shade, reflect, and refract using the



Q: How do we interpolate to get N? (Don't forget to normalize!!)

Epsilons

Due to finite precision arithmetic, we do not always get the exact intersection at a surface.

Q: What kinds of problems might this cause?

Q: How might we resolve this?

Intersecting with xformed geometry Intersecting with xformed geometry In general, objects will be placed using transformations. What if the object being intersected were transformed by a matrix M? The intersected normal is in object (local) coordinates. How do we transform it to world coordinates? Apply M¹ to the ray first and intersect in object (local) coordinates! 1