

Ray Tracing

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Reading

Required:

- ♦ Shirley, section 10.1-10.7 (handout)
- ♦ Triangle intersection handout

Further reading:

- ♦ Shirley errata on syllabus page, needed if you work from his book instead of the handout, which has already been corrected.
- ♦ T. Whitted. An improved illumination model for shaded display. Communications of the ACM 23(6), 343-349, 1980.
- ♦ A. Glassner. An Introduction to Ray Tracing. Academic Press, 1989.
- ♦ K. Turkowski, "Properties of Surface Normal Transformations," Graphics Gems, 1990, pp. 539-547.

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Geometric optics

Modern theories of light treat it as both a wave and a particle.

We will take a combined and somewhat simpler view of light – the view of **geometric optics**.

Here are the rules of geometric optics:

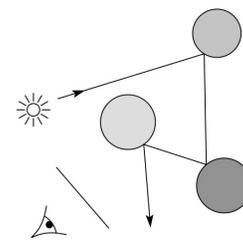
- ♦ Light is a flow of photons with wavelengths. We'll call these flows "light rays."
- ♦ Light rays travel in straight lines in free space.
- ♦ Light rays do not interfere with each other as they cross.
- ♦ Light rays obey the laws of reflection and refraction.
- ♦ Light rays travel from the light sources to the eye, but the physics is invariant under path reversal (reciprocity).

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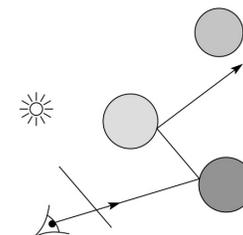
Eye vs. light ray tracing

Where does light begin?

At the light: light ray tracing (a.k.a., forward ray tracing or photon tracing)



At the eye: eye ray tracing (a.k.a., backward ray tracing)



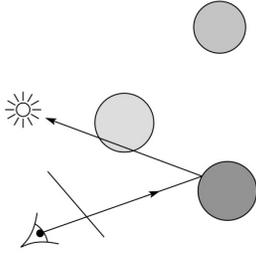
We will generally follow rays from the eye into the scene.

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Precursors to ray tracing

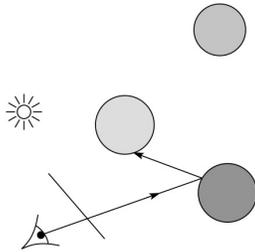
Local illumination

- ◆ Cast one eye ray, then shade according to light



Appel (1968)

- ◆ Cast one eye ray + one ray to light

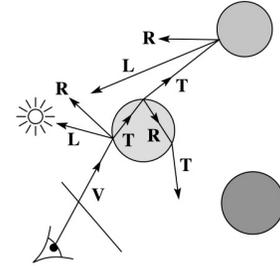


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Whitted ray-tracing algorithm

In 1980, Turner Whitted introduced ray tracing to the graphics community.

- ◆ Combines eye ray tracing + rays to light
- ◆ Recursively traces rays



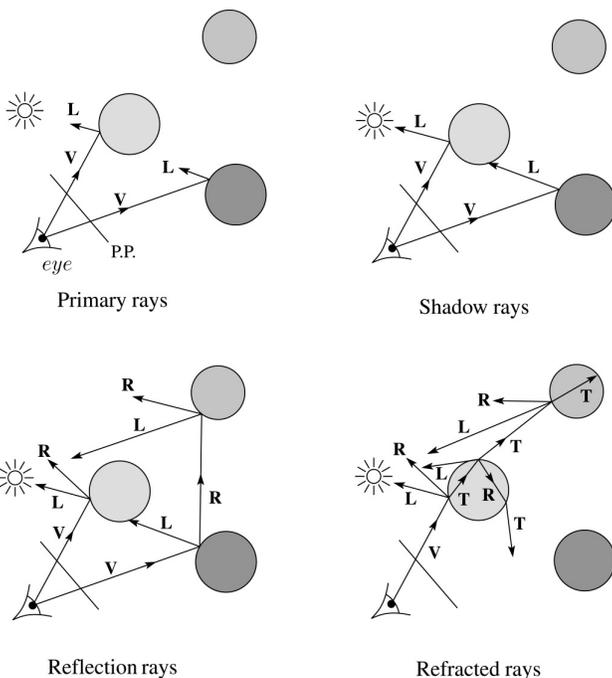
Algorithm:

1. For each pixel, trace a **primary ray** in direction V to the first visible surface.
2. For each intersection, trace **secondary rays**:
 - ◆ **Shadow rays** in directions L_i to light sources
 - ◆ **Reflected ray** in direction R .
 - ◆ **Refracted ray or transmitted ray** in direction T .

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Whitted algorithm (cont'd)

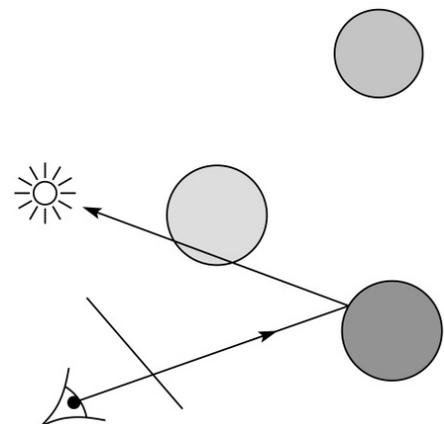
Let's look at this in stages:



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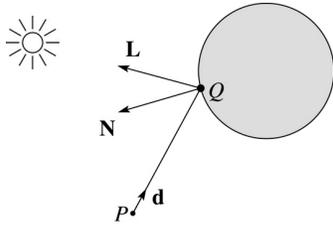
Ray casting and local illumination

Now let's actually build the ray tracer in stages. We'll start with ray casting and local illumination:



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Direct illumination



A ray is defined by an origin \mathbf{P} and a unit direction \mathbf{d} and is parameterized by $t > 0$:

$$\mathbf{P} + t\mathbf{d}$$

Let $I(\mathbf{P}, \mathbf{d})$ be the intensity seen along a ray. Then:

$$I(\mathbf{P}, \mathbf{d}) = I_{\text{direct}}$$

where

- I_{direct} is computed from the Phong model

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Ray-tracing pseudocode

We build a ray traced image by casting rays through each of the pixels.

function *tracelImage* (scene):

for each pixel (i,j) in image

$A = \text{pixelToWorld}(i,j)$

$P = \text{COP}$

$\mathbf{d} = (A - P) / \|A - P\|$

$I(i,j) = \text{traceRay}(\text{scene}, P, \mathbf{d})$

end for

end function

function *traceRay*(scene, P, \mathbf{d}):

$(t, \mathbf{N}, \text{mtrl}) \leftarrow \text{scene.intersect}(P, \mathbf{d})$

$Q \leftarrow \text{ray}(P, \mathbf{d})$ evaluated at t

$I = \text{shade}(\quad)$

return I

end function

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Shading pseudocode

Next, we need to calculate the color returned by the *shade* function.

function *shade*(mtrl, scene, $Q, \mathbf{N}, \mathbf{d}$):

$I \leftarrow \text{mtrl}.k_e + \text{mtrl}.k_a * I_a$

for each light source L **do**:

$\text{atten} = L \rightarrow \text{distanceAttenuation}(\quad)$

$I \leftarrow I + \text{atten} * (\text{diffuse term} + \text{specular term})$

end for

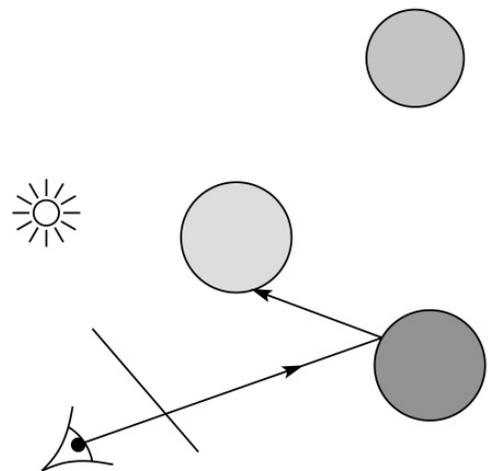
return I

end function

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Ray casting with shadows

Now we'll add shadows by casting shadow rays:



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Shading with shadows

To include shadows, we need to modify the shade function:

```

function shade(mtrl, scene, Q, N, d):
    I ← mtrl.ke + mtrl.ka * Ia
    for each light source L do:
        atten = L -> distanceAttenuation(Q) *
            L -> shadowAttenuation(      )
        I ← I + atten*(diffuse term + specular term)
    end for
    return I
end function
  
```

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Shadow attenuation

Computing a shadow can be as simple as checking to see if a ray makes it to the light source.

For a point light source:

```

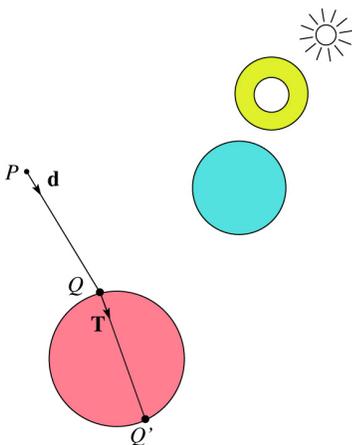
function PointLight::shadowAttenuation(scene, P)
    d = (this.position - P).normalize()
    (t, N, mtrl) ← scene.intersect(P, d)
    Compute tlight
    if (t < tlight) then:
        atten = 0
    else
        atten = 1
    end if
    return atten
end function
  
```

For a directional light, $t_{\text{light}} = \infty$.

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Shadow attenuation (cont'd)

Q: What if there are transparent objects along a path to the light source?



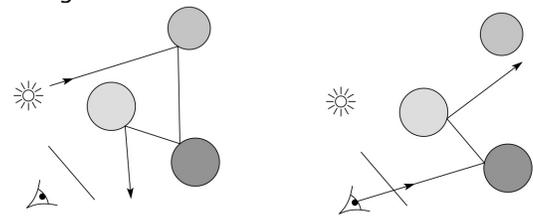
Suppose for simplicity that each object has a multiplicative transparency constant, k_t , which gets factored in every time an object is entered, possibly more than once for the same object.

Shirley's textbook describes a better attenuation model based on Beer's Law, which you can implement for extra credit.

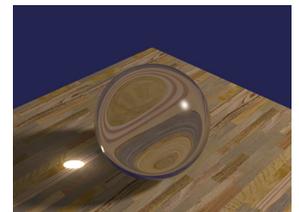
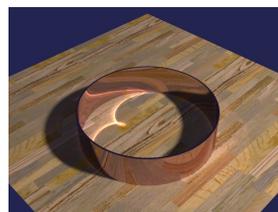
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Photon mapping

Combine light ray tracing (photon tracing) and eye ray tracing:



...to get **photon mapping**.



Renderings by Henrik Wann Jensen:
<http://graphics.ucsd.edu/~henrik/images/caustics.html>

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Shading in "Trace"

The Trace project uses a version of the Phong shading equation we derived in class, with two modifications:

- Distance attenuation is clamped to be at most 1:

$$A_j^{dist} = \min \left\{ 1, \frac{1}{a_j + b_j d_j + c_j d_j^2} \right\}$$

- Shadow attenuation A^{shadow} is included.

Here's what it should look like:

$$I = k_e + k_a L_a + \sum_j A_j^{shadow} A_j^{dist} L_j \left[k_d (\mathbf{N} \cdot \mathbf{L}_j)_+ + k_s (\mathbf{V} \cdot \mathbf{R}_j)_+^{n_s} \right]$$

I.e., we are not using the OpenGL shading equation, which is somewhat different.

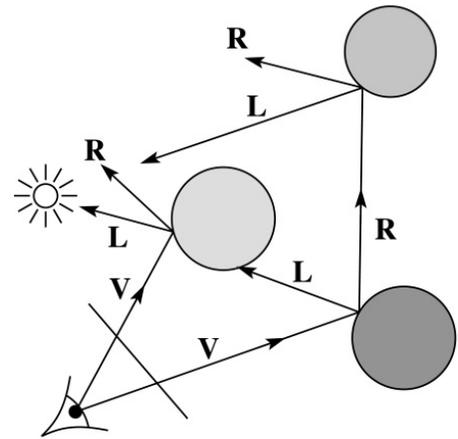
Note: the "R" here is the reflection of the *light* about the surface normal.

You must use the shading equation on this slide, not the one in Shirley's textbook.

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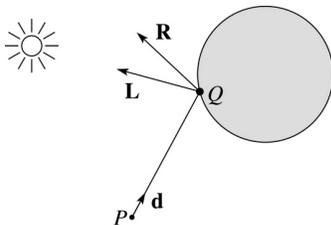
Recursive ray tracing with reflection

Now we'll add reflection:



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Shading with reflection



Let $I(P, \mathbf{d})$ be the intensity seen along a ray. Then:

$$I(P, \mathbf{d}) = I_{\text{direct}} + I_{\text{reflected}}$$

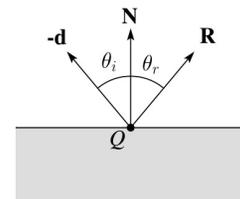
where

- I_{direct} is computed from the Phong model, plus shadow attenuation
- $I_{\text{reflected}} = k_r I(Q, \mathbf{R})$

Typically, we set $k_r = k_s$.

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Reflection



Law of reflection:

$$\theta_i = \theta_r$$

R is co-planar with **d** and **N**.

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Ray-tracing pseudocode, revisited

```

function traceRay(scene, P, d):
    (t, N, mtrl) ← scene.intersect(P, d)
    Q ← ray(P, d) evaluated at t
    I = shade(scene, mtrl, Q, N, -d)
    R = reflectDirection(          )
    I ← I + mtrl.kr * traceRay(scene, Q, R)
    return I
end function

```

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Terminating recursion

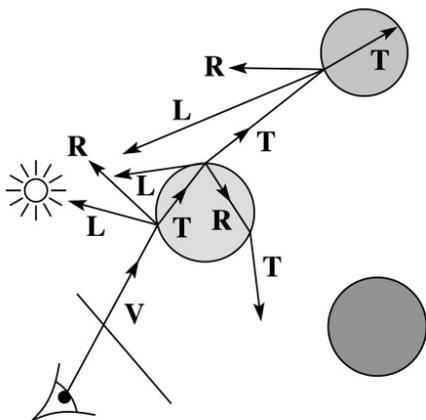
Q: How do you bottom out of recursive ray tracing?

Possibilities:

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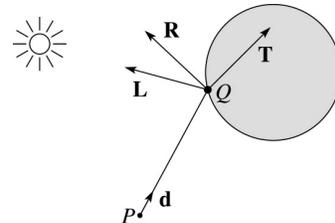
Whitted ray tracing

Finally, we'll add refraction, giving us the Whitted ray tracing model:



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Shading with reflection and refraction



Let $I(P, \mathbf{d})$ be the intensity seen along a ray. Then:

$$I(P, \mathbf{d}) = I_{\text{direct}} + I_{\text{reflected}} + I_{\text{transmitted}}$$

where

- ♦ I_{direct} is computed from the Phong model
- ♦ $I_{\text{reflected}} = k_r I(Q, \mathbf{R})$
- ♦ $I_{\text{transmitted}} = k_t I(Q, \mathbf{T})$

Typically, we set $k_r = k_s$ and $k_t = 1 - k_s$ (or 0, if opaque).

[Generally, k_r and k_t are determined by "Fresnel reflection," which depends on angle of incidence and changes the polarization of the light. This is discussed in Shirley's textbook and can be implemented for extra credit.]

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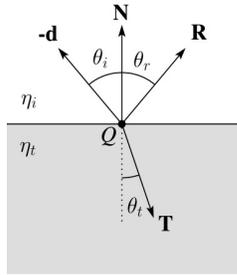
Refraction

Snell's law of refraction:

$$n_i \sin \theta_i = n_t \sin \theta_t$$

where n_i , n_t are **indices of refraction**.

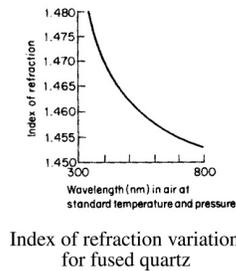
In all cases, **R** and **T** are co-planar with **d** and **N**.



The index of refraction is material dependent.

It can also vary with wavelength, an effect called **dispersion** that explains the colorful light rainbows from prisms. (We will generally assume no dispersion.)

Medium	Index of refraction
Vacuum	1
Air	1.0003
Water	1.33
Fused quartz	1.46
Glass, crown	1.52
Glass, dense flint	1.66
Diamond	2.42



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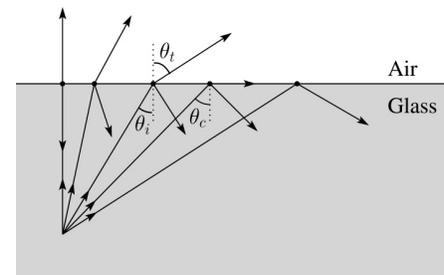
Total Internal Reflection

The equation for the angle of refraction can be computed from Snell's law:

What happens when $n_i > n_t$?

When θ_i is exactly 90° , we say that θ_i has achieved the "critical angle" θ_c .

For $\theta_i > \theta_c$, *no rays are transmitted*, and only reflection occurs, a phenomenon known as "total internal reflection" or TIR.



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Shirley handout

Watt uses different symbols. Here is the translation between them:

$$\begin{aligned} \mathbf{r} &= \mathbf{R} \\ \mathbf{t} &= \mathbf{T} \\ \phi &= \theta_t \\ \theta &= \theta_r = \theta_i \\ n &= n_i \\ n_t &= n_t \end{aligned}$$

Also, Shirley has two important errors that have already been corrected in the handout.

But, if you're consulting the original text, be sure to refer to the errata posted on the syllabus and on the project page for corrections.

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Ray-tracing pseudocode, revisited

```
function traceRay(scene, P, d):
    (t, N, mtrl) ← scene.intersect(P, d)
    Q ← ray(P, d) evaluated at t
    I = shade(scene, mtrl, Q, N, -d)
    R = reflectDirection(N, -d)
    I ← I + mtrl.kr * traceRay(scene, Q, R)
    if ray is entering object then
        n_i = index_of_air
        n_t = mtrl.index
    else
        n_i = mtrl.index
        n_t = index_of_air
    if (not TIR (
        T = refractDirection (
        I ← I + mtrl.kt * traceRay(scene, Q, T)
    end if
    return I
end function
```

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Terminating recursion, incl. refraction

Q: Now how do you bottom out of recursive ray tracing?

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Normals and shading when inside

When a ray is inside an object and intersects the object's surface on the way out, the normal will be pointing **away** from the ray (i.e., the normal always points to the outside by default).

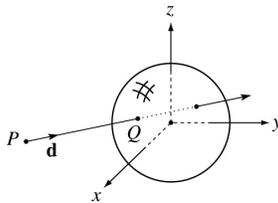
You must **negate** the normal before doing any of the shading, reflection, and refraction that follows.

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Intersecting rays with spheres

Now we've done everything except figure out what that "scene.intersect(P , \mathbf{d})" function does.

Mostly, it calls each object to find out the t value at which the ray intersects the object. Let's start with intersecting spheres...



Given:

- The coordinates of a point along a ray passing through P in the direction \mathbf{d} are:

$$x = P_x + td_x$$

$$y = P_y + td_y$$

$$z = P_z + td_z$$

- A unit sphere S centered at the origin defined by the equation:

Find: The t at which the ray intersects S .

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Intersecting rays with spheres

Solution by substitution:

$$x^2 + y^2 + z^2 - 1 = 0$$

$$(P_x + td_x)^2 + (P_y + td_y)^2 + (P_z + td_z)^2 - 1 = 0$$

$$at^2 + bt + c = 0$$

where

$$a = d_x^2 + d_y^2 + d_z^2$$

$$b = 2(P_x d_x + P_y d_y + P_z d_z)$$

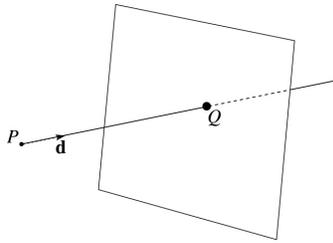
$$c = P_x^2 + P_y^2 + P_z^2 - 1$$

Q: What are the solutions of the quadratic equation in t and what do they mean?

Q: What is the normal to the sphere at a point (x, y, z) on the sphere?

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Ray-plane intersection



We can write the equation of a plane as:

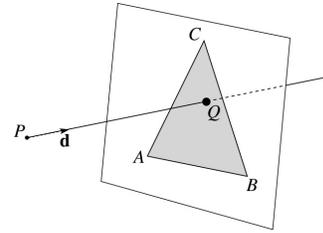
$$ax + by + cz = d$$

The coefficients a , b , and c form a vector that is normal to the plane, $\mathbf{n} = [a \ b \ c]^T$. Thus, we can re-write the plane equation as:

We can solve for the intersection parameter (and thus the point):

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Ray-triangle intersection



To intersect with a triangle, we first solve for the equation of its supporting plane.

How might we compute the (un-normalized) normal?

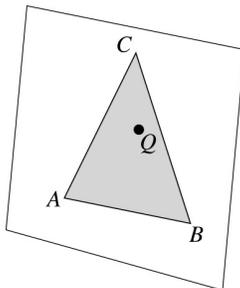
Given this normal, how would we compute d ?

Using these coefficients, we can solve for Q . Now, we need to decide if Q is inside or outside of the triangle.

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3D inside-outside test

One way to do this “inside-outside test,” is to see if Q lies on the left side of each edge as we move counterclockwise around the triangle.

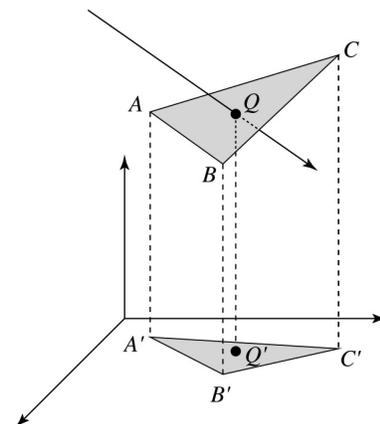


How might we use cross products to do this?

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2D inside-outside test

Without loss of generality, we can perform this same test after projecting down a dimension:



If Q' is inside of $A'B'C'$, then Q is inside of ABC .

Why is this projection desirable?

Which axis should you “project away”?

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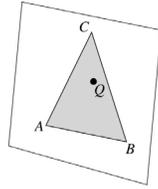
Barycentric coordinates

As we'll see in a moment, it is often useful to represent Q as an **affine combination** of A , B , and C :

$$Q = \alpha A + \beta B + \gamma C$$

where:

$$\alpha + \beta + \gamma = 1$$



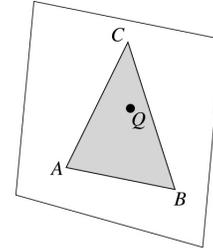
We call α , β , and γ the **barycentric coordinates** of Q with respect to A , B , and C .

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Computing barycentric coordinates

Given a point Q that is inside of triangle ABC , we can solve for Q 's barycentric coordinates in a simple way:

$$\alpha = \frac{\text{Area}(QBC)}{\text{Area}(ABC)} \quad \beta = \frac{\text{Area}(AQC)}{\text{Area}(ABC)} \quad \gamma = \frac{\text{Area}(ABQ)}{\text{Area}(ABC)}$$



How can cross products help here?

In the end, these calculations can be performed in the 2D projection as well!

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Epsilons

Due to finite precision arithmetic, we do not always get the exact intersection at a surface.

Q: What kinds of problems might this cause?

Q: How might we resolve this?

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Summary

What to take home from this lecture:

- ♦ The meanings of all the boldfaced terms.
- ♦ Enough to implement basic recursive ray tracing.
- ♦ How reflection and transmission directions are computed.
- ♦ How ray--object intersection tests are performed on spheres, planes, and triangles
- ♦ How barycentric coordinates within triangles are computed
- ♦ How ray epsilons are used.

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Interpolating vertex properties

The barycentric coordinates can also be used to interpolate vertex properties such as:

- ◆ material properties
- ◆ texture coordinates
- ◆ normals

For example:

$$k_d(Q) = \alpha k_d(A) + \beta k_d(B) + \gamma k_d(C)$$

Interpolating normals, known as Phong interpolation, gives triangle meshes a smooth shading appearance. (Note: don't forget to normalize interpolated normals.)