## Computing light source directions

Trick: place a chrome sphere in the scene


- the location of the highlight tells you where the light source is


## Recall the rule for specular reflection

For a perfect mirror, light is reflected about $\mathbf{N}$


$$
I_{e}=\left\{\begin{array}{cl}
I_{i} & \text { if } \mathbf{V}=\mathbf{R} \\
0 & \text { otherwise }
\end{array}\right.
$$

We see a highlight when $\mathbf{V}=\mathbf{R}$

- then $L$ is given as follows:

$$
\mathbf{L}=2(\mathbf{N} \cdot \mathbf{R}) \mathbf{N}-\mathbf{R}
$$

## Computing the light source direction

Chrome sphere that has a highlight at position $\mathbf{h}$ in the image

image plane
Can compute $\theta$ (and hence $\mathbf{N}$ ) from this figure Now just reflect $\mathbf{V}$ about $\mathbf{N}$ to obtain $\mathbf{L}$

## Depth from normals



What we have


What we want

## Depth from normals



Get a similar equation for $\mathbf{V}_{\mathbf{2}}$

- Each normal gives us two linear constraints on z
- compute $z$ values by solving a matrix equation
- On the boundary we have only one constraint.


## Trick for handling shadows

Weight each equation by the pixel brightness:

$$
I_{i}\left(I_{i}\right)=I_{i}\left[k_{d} \mathbf{N} \cdot \mathbf{L}_{\mathbf{i}}\right]
$$

Gives weighted least-squares matrix equation:

$$
\left[\begin{array}{lll}
I_{1}^{2} & \ldots & I_{n}^{2}
\end{array}\right]=k_{d} \mathbf{N}^{T}\left[\begin{array}{lll}
I_{1} \mathbf{L}_{1} & \ldots & I_{n} \mathbf{L}_{\mathbf{n}}
\end{array}\right]
$$

Solve for $\mathrm{N}, \mathrm{k}_{\mathrm{d}}$ as before

## Example



## Results...


from Athos Georghiades
http://cvc.yale.edu/people/Athos.html

## Limitations

- doesn't work for shiny things, semi-translucent things
- shadows, inter-reflections are difficult
- Single light source illumination
- camera and lights have to be distant
- calibration requirements
- measure light source directions, intensities
- camera response function


## Newer work addresses some of these issues

Some pointers for further reading:

- Zickler, Belhumeur, and Kriegman, "Helmholtz Stereopsis: Exploiting Reciprocity for Surface Reconstruction." IJCV, Vol. 49 No. 2/3, pp 215-227.
- Hertzmann \& Seitz, Example-Based Photometric Stereo: Shape Reconstruction with General, Varying BRDFs." IEEE Trans. PAMI 2005
- Basri, Jacobs and Kemelmacher Photometric Stereo with General Unknown Lighting, International Journal of Computer Vision (IJCV) 2007

Hertzmann \& Seitz, Example-Based Photometric Stereo: Shape Reconstruction with General, Varying BRDFs." IEEE Trans. PAMI 2005

## Shiny things










स思：：


## Virtual views



## Velvet



## Virtual Views

## Brushed Fur



## Brushed Fur



## Virtual Views



- Basri, Jacobs and Kemelmacher Photometric Stereo with General Unknown Lighting, International Journal of Computer Vision (IJCV) 2007


## Illumination Modeling





## Spherical Harmonics



## Cumulative Energy



## Photometric Stereo

Factorization to Shape and Light


SVD recovers $L$ and $S$ up to an ambiguity

## Results with arbitrary unknown lighting



## 3D Face Shape Reconstruction

Input: Single Image, general lighting, pose


Output: 3D shape


Kemelmacher and Basri Molding Face Shapes by Example IEEE Trans. PAMI

## One approach

- Represent input face as a combination of hundreds of 3D face in a database:



## Our approach

> Input image Single reference model


## Harmonic representation of lighting

| albedo <br> n <br> surface normal <br> $\mathrm{n}=(\mathrm{n}, n$ |
| :--- |
|  |
| Positive values <br>  <br> Negative values |

$$
\begin{array}{r}
\mathbb{I}(x, y)=\mathbb{R}(\mathbf{n} ; \rho, \mathbf{l}) \approx \rho \sum_{\mathrm{i}=\mathbf{0}}^{\mathrm{K}-1} 1_{\mathrm{i}} Y_{\mathrm{i}}(\mathbf{n}) \\
\\
37.6 \%
\end{array}
$$


87.5\%

In practice
~95\% accuracy

- Lambertian reflectance
- Accounts for attached shadows
- Arbitrary illumination (point sources,

99.2\% diffuse, combinations...)


## Formulation

$$
\mathrm{I}(\mathrm{x}, \mathrm{y})=\mathrm{R}(\mathbf{n} ; \rho, \mathbf{l}) \approx \rho \sum_{\mathrm{i}=\mathbf{0}}^{\mathrm{K}-\mathbf{1}} 1_{\mathrm{i}} Y_{\mathrm{i}}(\mathbf{n})
$$

Unknown surface normal:
$\mathbf{n}(\mathrm{x}, \mathrm{y})=\frac{\mathbf{1}}{\sqrt{\mathrm{p}^{2}+\mathrm{q}^{2}+\mathbf{1}}}(\mathrm{p}, \mathrm{q},-\mathbf{1})^{\mathrm{T}}$
$p(x, y)=\partial z / \partial x$
$q(x, y)=\partial z / \partial y$

Known image:


Known reference model:

$\min _{1, z, \rho} \int_{\Omega}\left(I(x, y)-\rho \sum_{i=0}^{K-1} 1_{i} Y_{i}(n)\right)^{2}+\lambda_{1}\left(\Delta G * d_{z}(x, y)\right)^{2}+\lambda_{2}\left(\Delta G * d_{\rho}(x, y)\right)^{2} d x d y$

$$
d_{z}(x, y)=z(x, y)-z_{r e f}(x, y) \quad d_{\rho}(x, y)=\rho(x, y)-\rho_{\text {ref }}(x, y)
$$

## Reconstruction steps

$$
\min _{1, z, \rho} \int_{\Omega}\left(I(x, y)-\rho \sum_{i=0}^{K-1} 1_{i} Y_{i}(n)\right)^{2}+\lambda_{1}\left(\Delta G_{i} *_{z}(x, y)\right)^{2}+\lambda_{2}\left(\Delta G^{*} *_{\rho}(x, y)\right)^{2} d x d y
$$

Solve for lighting:


Different faces experience the same lighting in similar way. So use reference face to fill in the missing data.
Solve for depth:

$$
\begin{aligned}
& \mathrm{I}(\mathrm{x}, \mathrm{y})=\rho_{\mathrm{ref}} 1_{0}+\frac{\rho_{\text {ref }}}{\mathrm{N}_{\text {ref }}}\left(1_{1} \mathrm{p}+1_{2} \mathrm{q}-1_{3}\right) \\
& \mathrm{N}=\sqrt{\mathrm{p}^{2}+\mathrm{q}^{2}+1}
\end{aligned}
$$

Linear PDE in $z(x, y)$
Boundaries: We let partial derivatives across boundaries to vanish

## Results

## on images downloaded from the internet



## Expressions



## Results: unknown pose



