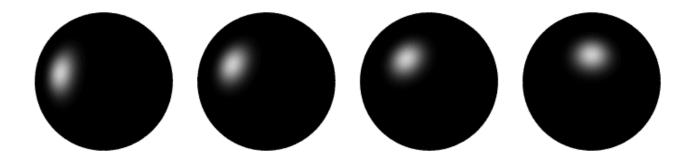
## Computing light source directions

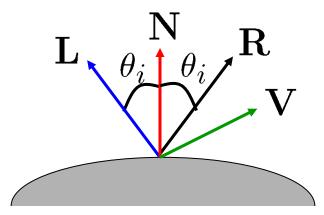
Trick: place a chrome sphere in the scene



· the location of the highlight tells you where the light source is

## Recall the rule for specular reflection

For a perfect mirror, light is reflected about N



$$I_e = \begin{cases} I_i & \text{if } \mathbf{V} = \mathbf{R} \\ 0 & \text{otherwise} \end{cases}$$

We see a highlight when V = R

• then L is given as follows:

$$L = 2(N \cdot R)N - R$$

# Computing the light source direction

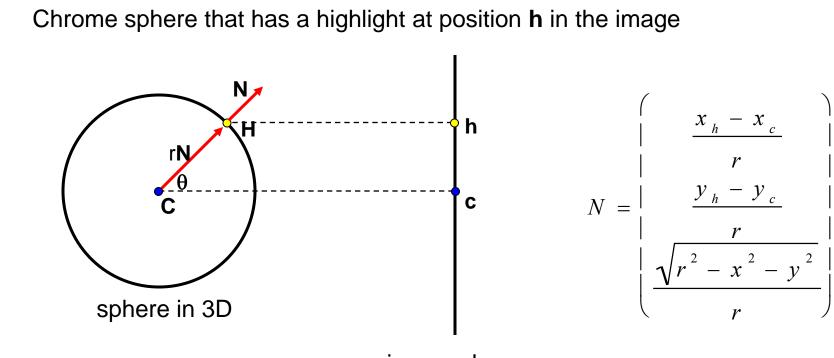
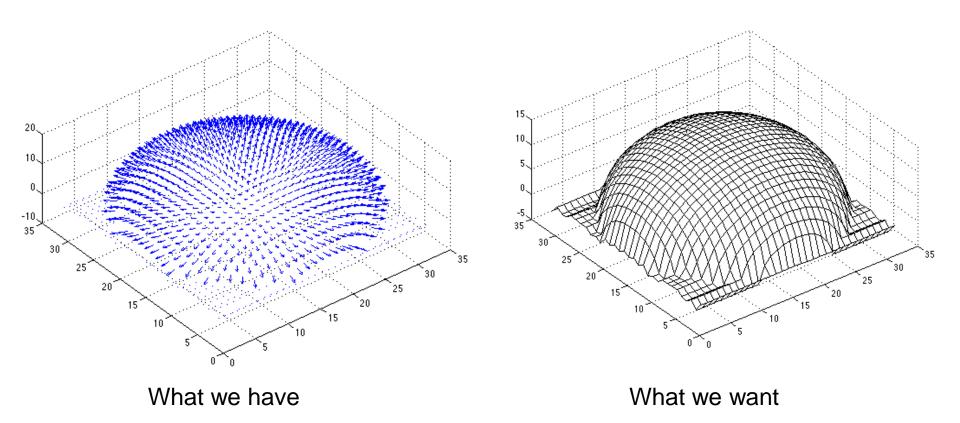


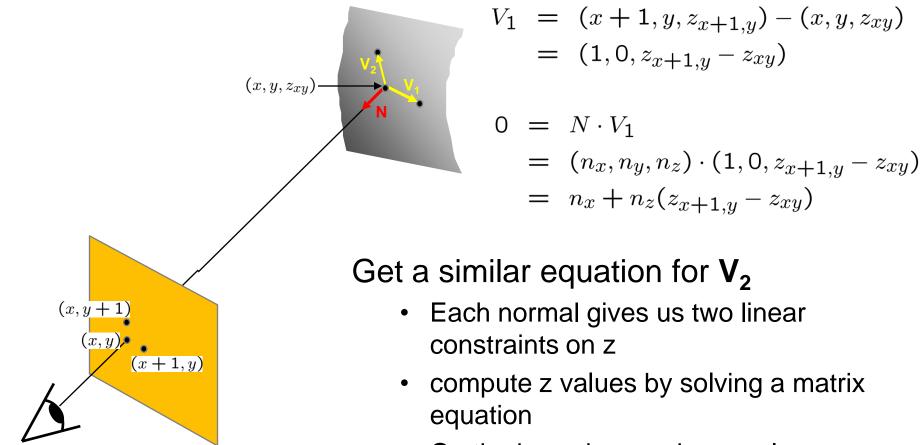
image plane

Can compute  $\theta$  (and hence **N**) from this figure Now just reflect **V** about **N** to obtain **L** 

## Depth from normals



## Depth from normals



• On the boundary we have only one constraint.

## Trick for handling shadows

Weight each equation by the pixel brightness:

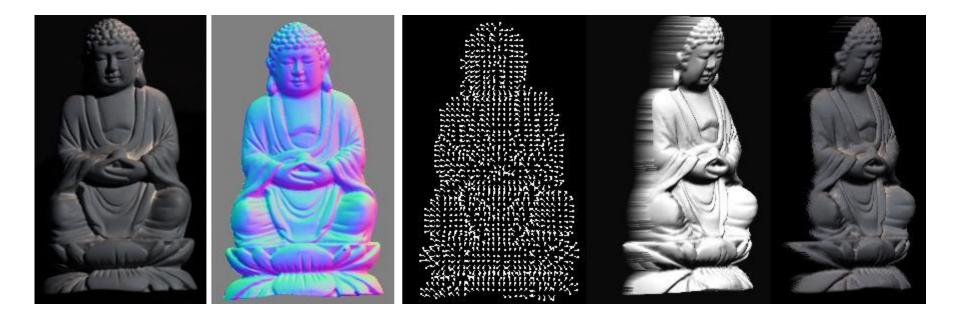
$$I_i(I_i) = I_i[k_d \mathbf{N} \cdot \mathbf{L_i}]$$

Gives weighted least-squares matrix equation:

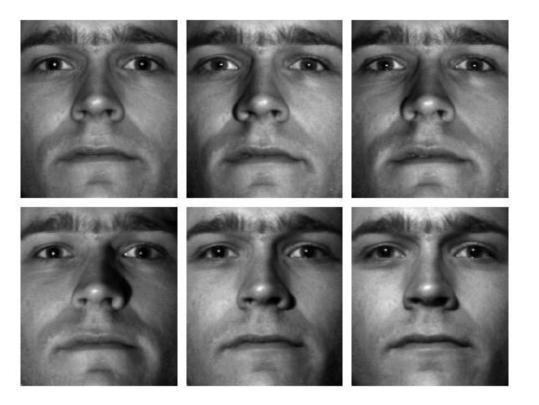
$$\begin{bmatrix} I_1^2 & \dots & I_n^2 \end{bmatrix} = k_d \mathbf{N}^T \begin{bmatrix} I_1 \mathbf{L}_1 & \dots & I_n \mathbf{L}_n \end{bmatrix}$$

Solve for N, k<sub>d</sub> as before

# Example



#### Results...





### from Athos Georghiades

http://cvc.yale.edu/people/Athos.html

## Limitations

- doesn't work for shiny things, semi-translucent things
- shadows, inter-reflections are difficult
- Single light source illumination
- camera and lights have to be distant
- calibration requirements
  - measure light source directions, intensities
  - camera response function

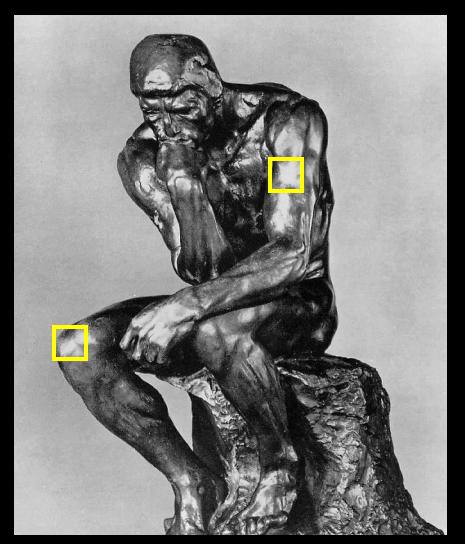
#### Newer work addresses some of these issues

#### Some pointers for further reading:

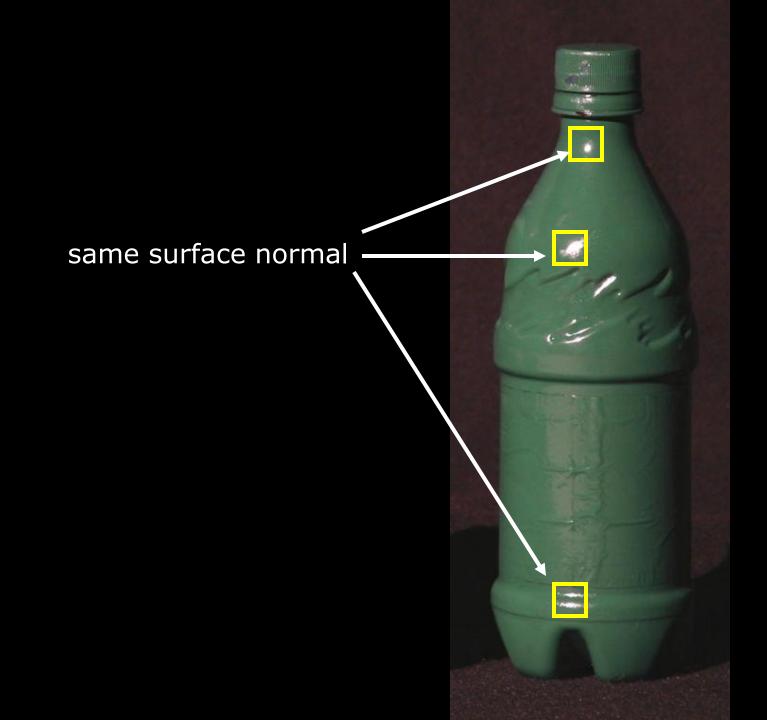
- Zickler, Belhumeur, and Kriegman, "<u>Helmholtz Stereopsis: Exploiting Reciprocity for</u> <u>Surface Reconstruction</u>." IJCV, Vol. 49 No. 2/3, pp 215-227.
- Hertzmann & Seitz, <u>Example-Based Photometric Stereo: Shape Reconstruction with</u> <u>General, Varying BRDFs</u>." IEEE Trans. PAMI 2005
- Basri, Jacobs and Kemelmacher <u>Photometric Stereo with General Unknown Lighting</u>, International Journal of Computer Vision (IJCV) 2007

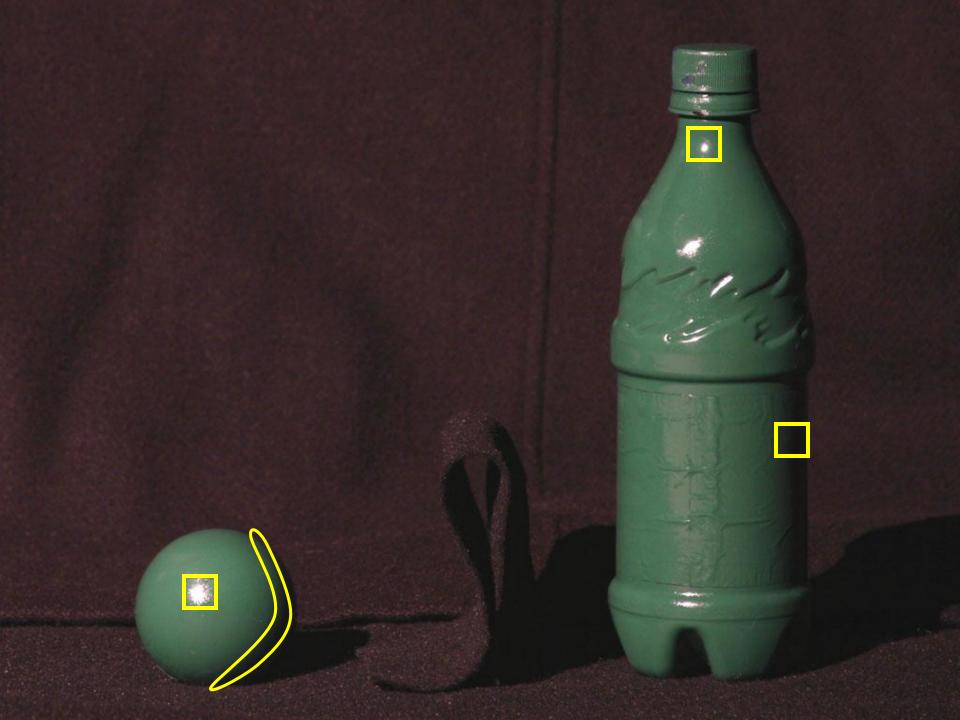
Hertzmann & Seitz, <u>Example-Based</u> <u>Photometric Stereo: Shape</u> <u>Reconstruction with General, Varying</u> <u>BRDFs</u>." IEEE Trans. PAMI 2005

# **Shiny things**



#### "Orientation consistency"







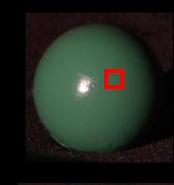




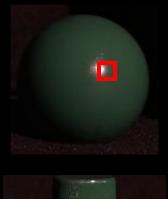
















# Virtual views









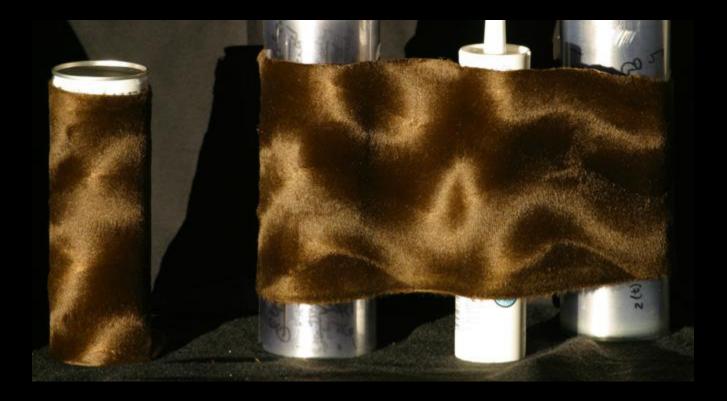
# **Virtual Views**



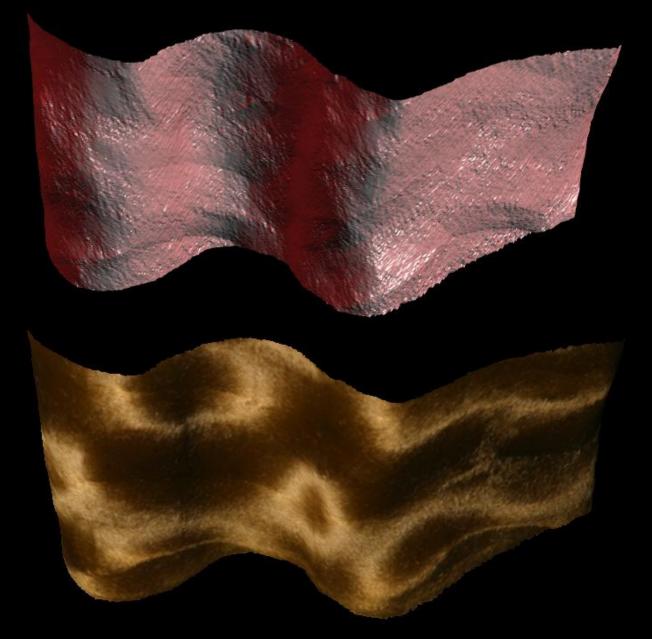
# **Brushed Fur**



# **Brushed Fur**

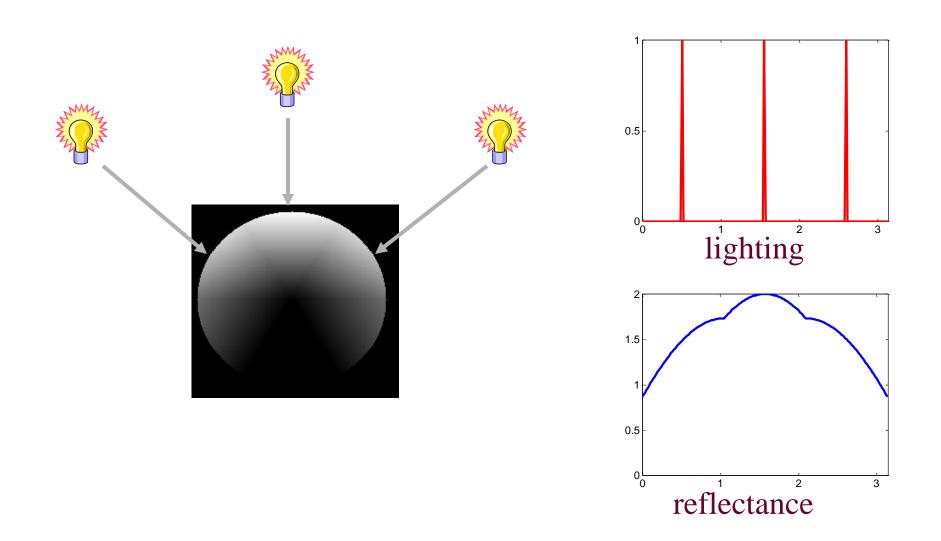


# **Virtual Views**

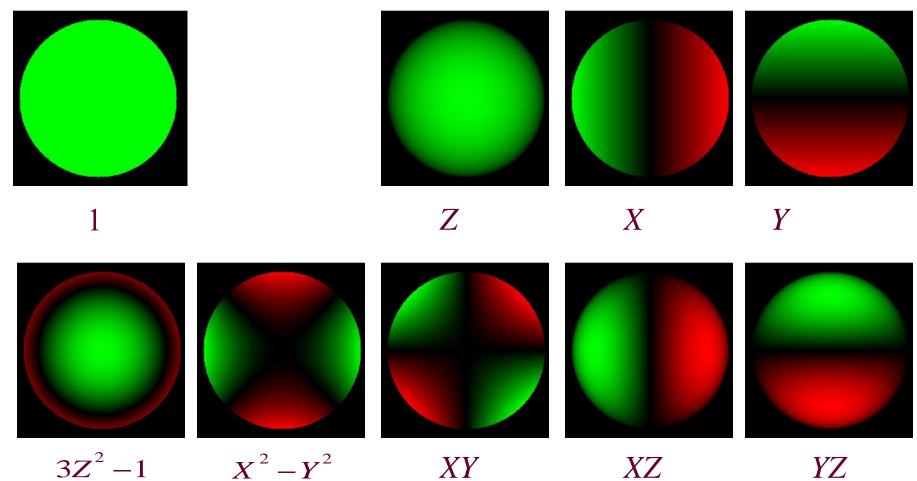


Basri, Jacobs and Kemelmacher
Photometric Stereo with General Unknown
Lighting, International Journal of Computer
Vision (IJCV) 2007

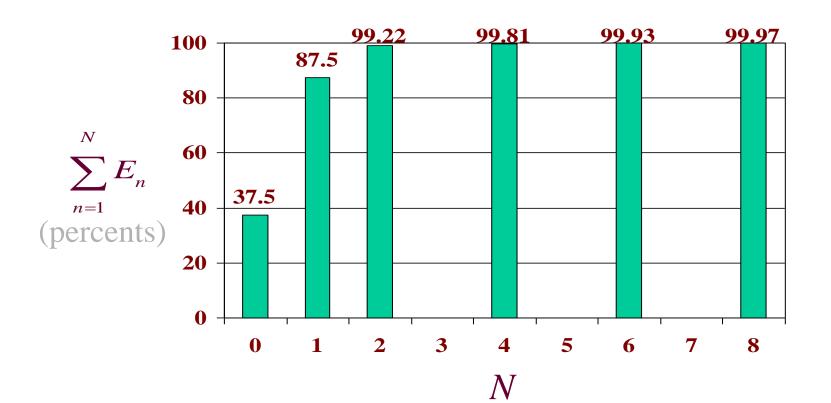
### **Illumination Modeling**



### **Spherical Harmonics**

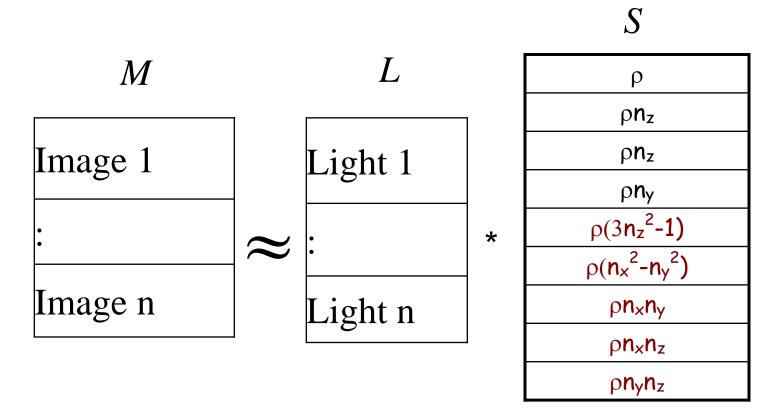


### **Cumulative Energy**



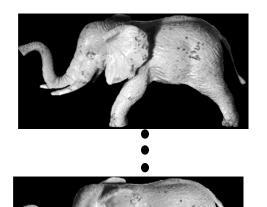
# Photometric Stereo

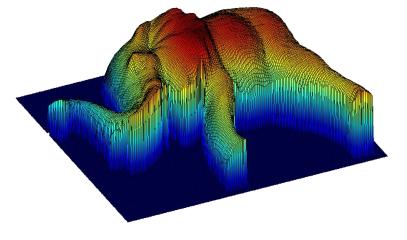
Factorization to Shape and Light

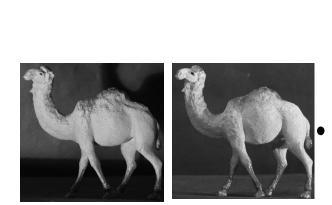


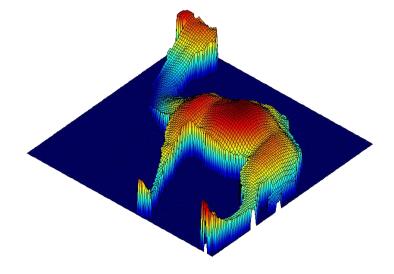
SVD recovers L and S up to an ambiguity

### Results with arbitrary unknown lighting





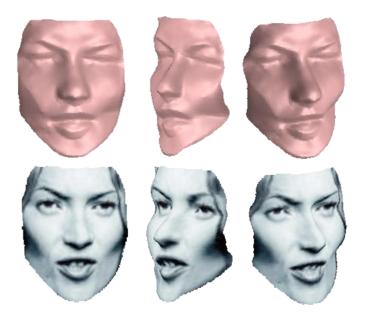




#### Input: Single Image, general lighting, pose



#### Output: 3D shape



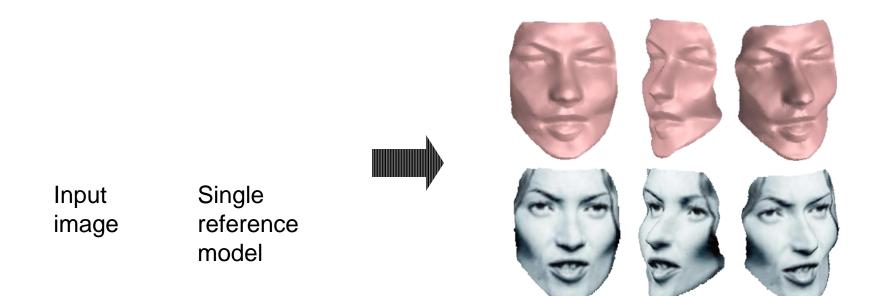
Kemelmacher and Basri Molding Face Shapes by Example IEEE Trans. PAMI

### One approach

 Represent input face as a combination of hundreds of 3D face in a database:



### Our approach

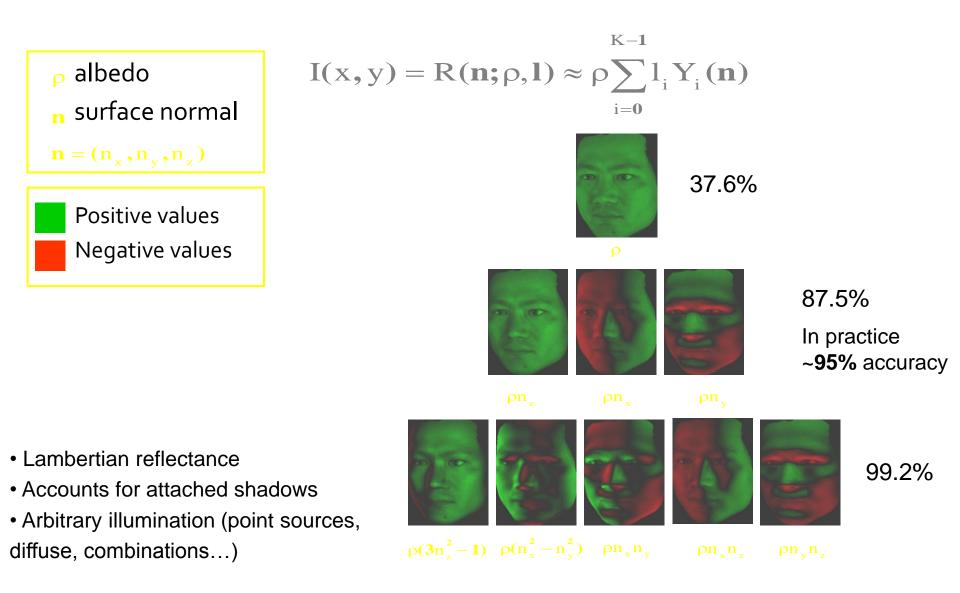




unknown lighting unknown boundaries unknown albedo significant differences in shape including

- expression
- gender
- race

# Harmonic representation of lighting



### Formulation

$$I(x, y) = R(n; \rho, l) \approx \rho \sum_{i=0}^{K-1} l_i Y_i(n)$$

Unknown surface normal:

$$\mathbf{n}(x,y) = \frac{1}{\sqrt{p^2 + q^2 + 1}} (p,q,-1)^T$$

 $p(x, y) = \partial z / \partial x$  $q(x, y) = \partial z / \partial y$ 

Unknown harmonic coeff.  $\mathbf{l} = (l_0 \dots l_{K-1})$ Unknown albedo  $\rho(\mathbf{x}, \mathbf{y})$  Known image:



Known reference model:



$$\min_{\mathbf{l},z,\rho} \iint_{\Omega} \left( I(x,y) - \rho \sum_{i=0}^{K-1} \mathbf{l}_i \mathbf{Y}_i(\mathbf{n}) \right)^2 + \lambda_1 (\Delta G * \mathbf{d}_z(x,y))^2 + \lambda_2 (\Delta G * \mathbf{d}_\rho(x,y))^2 dxdy$$

 $d_{z}(x,y) = z(x,y) - z_{ref}(x,y)$   $d_{\rho}(x,y) = \rho(x,y) - \rho_{ref}(x,y)$ 

# **Reconstruction steps**

$$\min_{\mathbf{l},z,\rho} \int_{\Omega} \left( I(\mathbf{x},\mathbf{y}) - \rho \sum_{i=0}^{K-1} l_i Y_i(\mathbf{n}) \right)^2 + \lambda_1 (\Delta G * d_z(\mathbf{x},\mathbf{y}))^2 + \lambda_2 (\Delta G * d_\rho(\mathbf{x},\mathbf{y}))^2 d\mathbf{x} d\mathbf{y}$$

#### Solve for lighting:

$$\min_{l} \sum_{(x,y)\in\Omega} \left( I(x,y) - \rho_{ref} \sum_{i=0}^{K-1} l_i Y_i(\mathbf{n}_{ref}) \right)^2$$

#### Solve for depth:

$$I(x, y) = \rho_{ref} l_0 + \frac{\rho_{ref}}{N_{ref}} (l_1 p + l_2 q - l_3)$$
$$N = \sqrt{p^2 + q^2 + 1}$$

Different faces experience the same lighting in similar way. So use reference face to fill in the missing data.

#### Linear PDE in z(x,y)

Boundaries: We let partial derivatives across boundaries to vanish

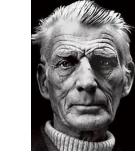
### Results on images downloaded from the internet











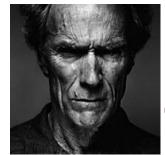






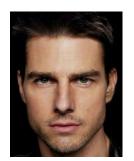












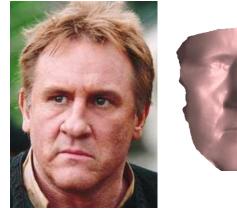






# Expressions



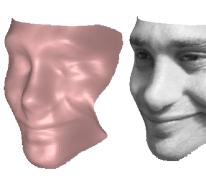




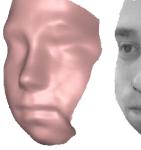






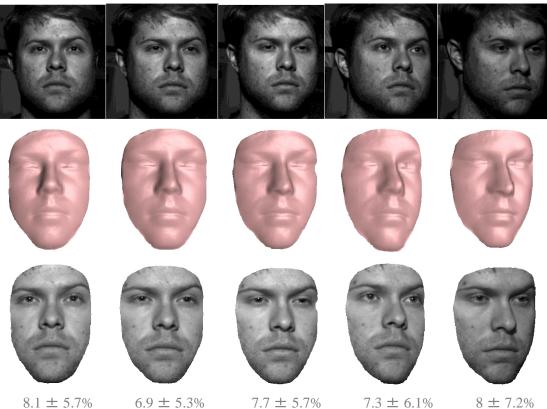








# Results: unknown pose



 $8.1 \pm 5.7\%$